

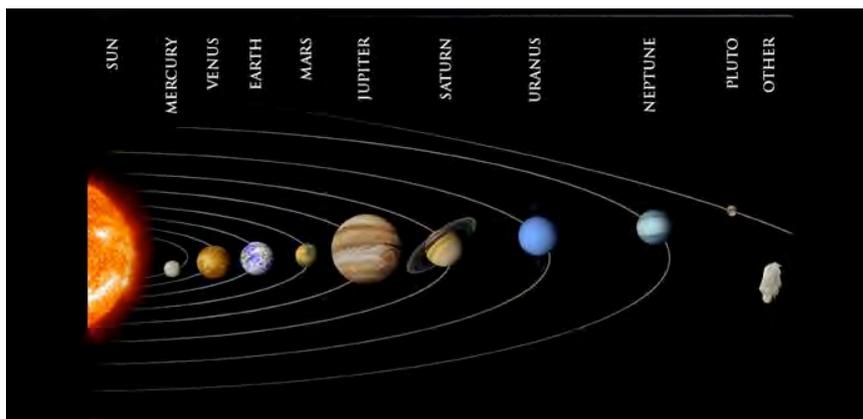
Recent Discoveries from the Analysis of Cassini Spacecraft Data in Saturn's Orbit

Part 1 : The Saturn system, its exploration by Voyager,
and the Cassini mission

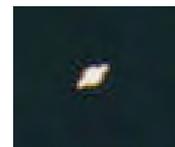
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France
philippe.zarka@obspm.fr

Invited Conference at Future University, Hakodate, Japon, 1/10/2004

- Saturn known from ancients
- 6th planet in solar system, farthest visible with naked eye

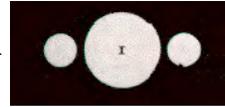


- Scientific exploration started in 17th century,
with first telescopic observations



• 1610

Galileo notes the “triple planet Saturn with his telescope
(1612 single, 1616 ansae)



• 1655

Huygens discovers Saturn's largest satellite, Titan



• 1659

Huygens proposes the concept of planetary rings
→ solid or particles ?



- 1671–84 Cassini discovers a division in the ring (→A&B)



- He also discovers the satellites Iapetus, Rhea, Dione and Tethys



Iapetus



Rhea

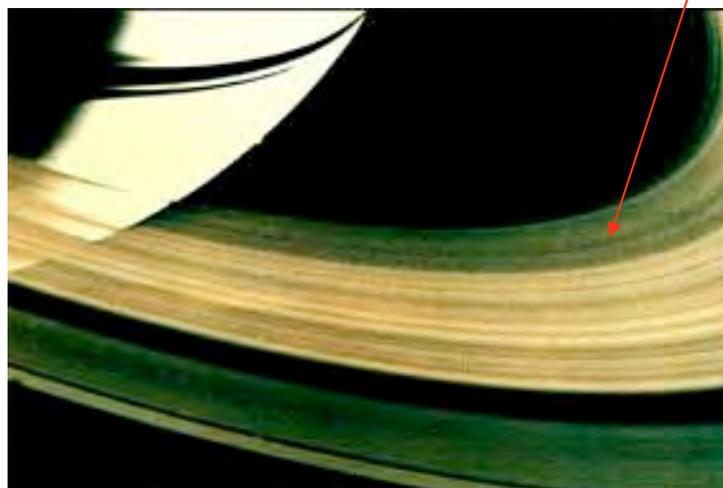


Dione



Tethys

- 1789 Herschel discovers satellites Mimas and Enceladus — and notes thinness of rings
- 1848 Bond and Lassell discover the satellite Hyperion
- 1850 Bond, Bond and Daws discover inner ring C

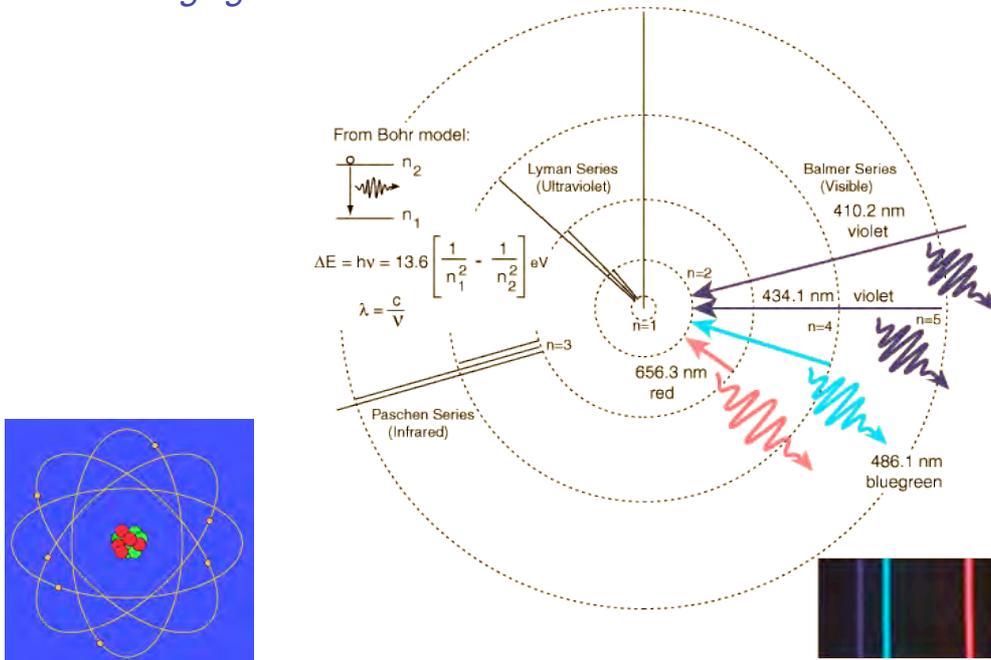


- 1857 Maxwell proves that rings are not solid (stability relative to Keplerian motion)

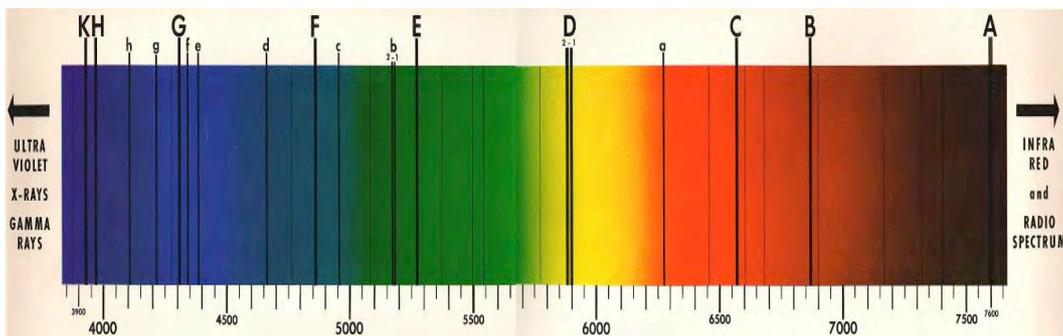


[SPECTROSCOPY]

Atoms and molecules change electrons' energy in absorbing or emitting light

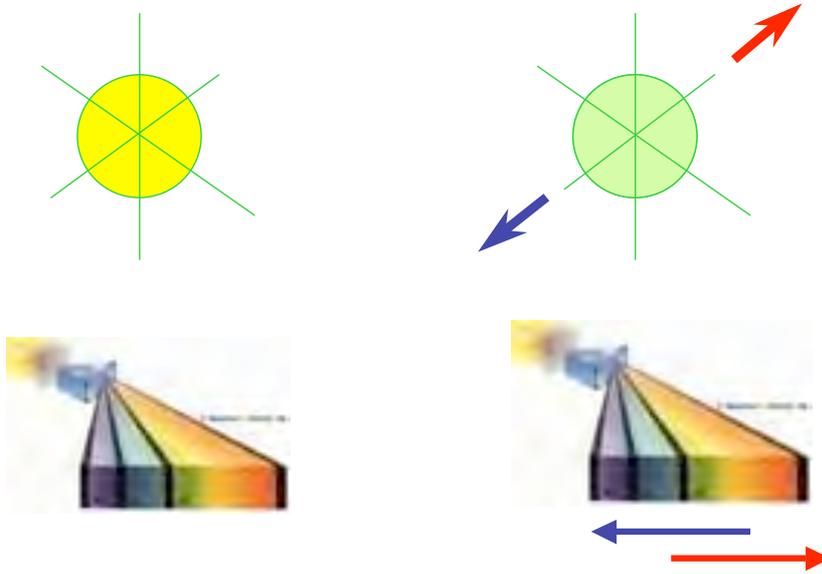


A series of lines is the fingerprint of a particular atom or molecule



→ SPECTROSCOPY reveals chemical composition

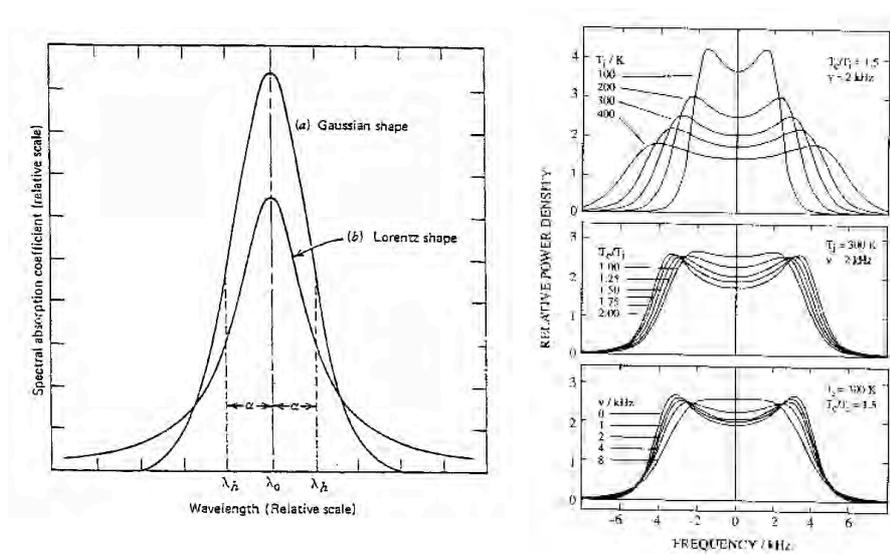
Relative motion source-observer shift the frequency of light ...



... or sound : DOPPLER effect



Shape of spectral line reveals temperature or rotation of source

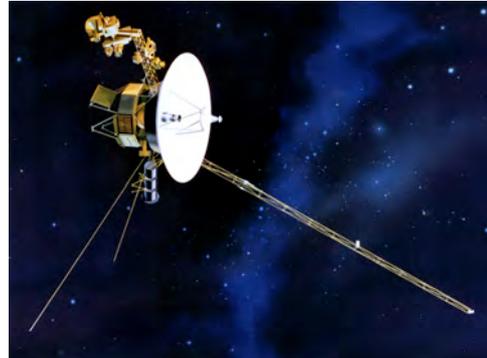


- 1872-95 Keeler measures ring velocities by spectroscopy
→ observational proof of ring particles
He discovers Keeler/Encke gap in A ring
- 1898 Pickering discovers satellite Phoebe
- 1932 Wildt discovers methane (CH₄) and ammonia (NH₃) on Saturn
- 1943–44 Kuiper discovers methane and ammonia on Titan



- 1957 Space exploration starts with Sputnik satellite
- 1960's Satellites orbiting Earth
- 1970-80's Planetary exploration : Mars, Venus, Mercury...

- 1979 Pioneer 11 flies past Saturn
- 1980 Voyager 1 encounters Saturn
- 1981 Voyager 2 encounters Saturn

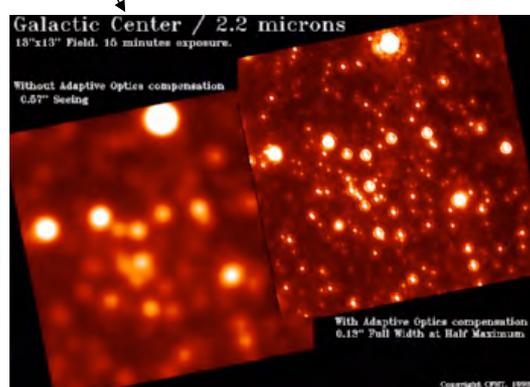


→ Next steps

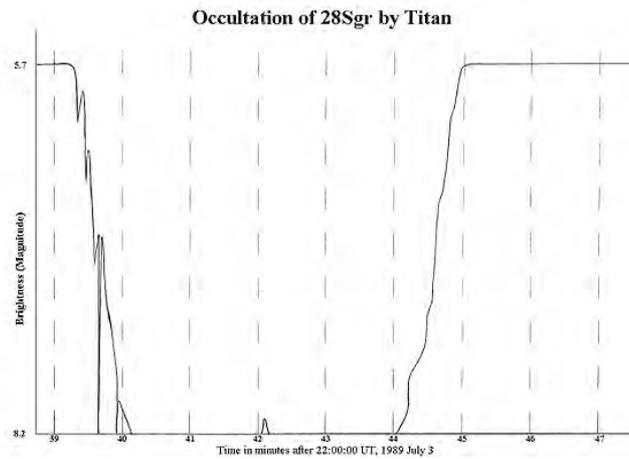
- Jupiter : Galileo orbital tour + probe (1995-2002)
- Saturn : Cassini orbital tour
+ Huygens Titan descent probe (2004-2008)

+ Remote observations :

- ≥ 1989 Hubble Space Telescope's Wide Field and Planetary Camera
- 1990's Adaptive Optics on ground-based IR telescopes



- Use of Star Occultations
(e.g. 28 Sgr observed from Earth in 1989)



+ Theory, Laboratory experiments, Numerical simulations

- Spacecraft instrumentation : Pioneer 11

Imaging Photopolarimeter (→ low-res images)

[MAG] Flux Gate Magnetometer

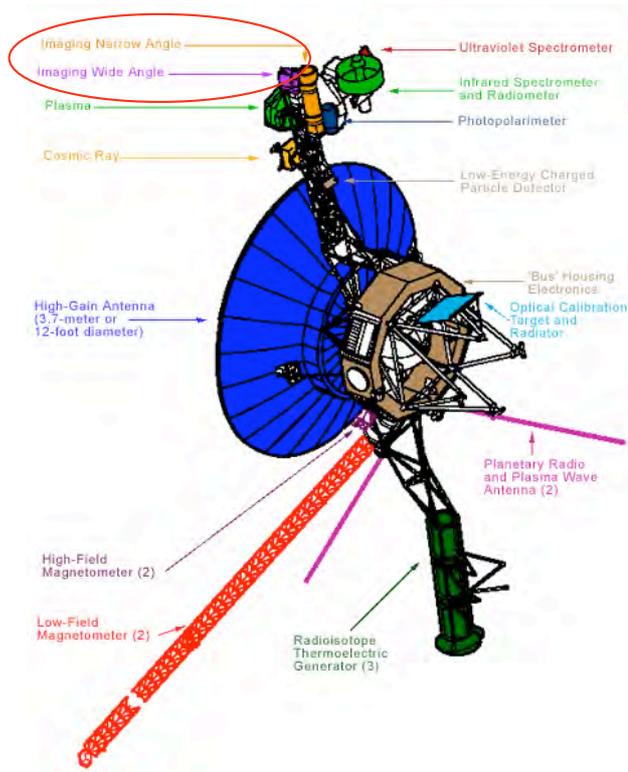
[PLS] Plasma Analyzer
Charged Particle Instrument
Trapped Radiation Detector

[UV] Ultraviolet Photometer

Cosmic Ray Telescope
Geiger Tube Telescope

(4 other experiments failed)

- Voyager 1-2 :
12 experiments
(0 failed)



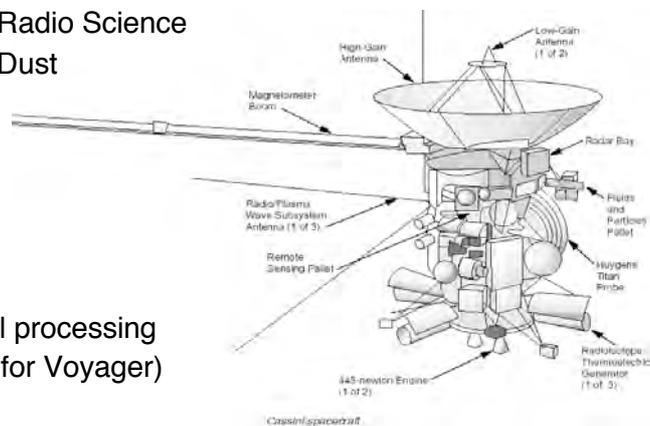
Cassini instrumentation :

12 experiments on Orbiter :

- [ISS] Imaging
- [UVIS, VIMS, CIRS] Spectroscopy
- [MAG] Magnetometer
- [CAPS, INMS, MIMI, **RPWS**] Waves & Particles
- [RSS] Radio Science
- [CDA] Dust
- [RADAR]

+ 6 experiments on Huygens probe

Extensive onboard signal processing
(400 microprocessors, 1 for Voyager)



SUMMARY OF
PRE-CASSINI KNOWLEDGE ABOUT SATURN
&
QUESTIONS TO BE ADDRESSED BY CASSINI

- (1) Saturn's interior**
- (2) Saturn's atmosphere**
- (3) Rings**
- (4) Titan**
- (5) Icy satellites**
- (6) Magnetosphere**

SATURN'S DATA

| | |
|-------------------|-----------------------|
| Orbital period : | 29.4 years |
| Semi-major axis : | 9.54 AU |
| Mass : | 95 x Earth's mass |
| Radius : | 9.4 x Earth's radius |
| Density : | 687 kg/m ³ |
| Composition : | Mostly H, He |

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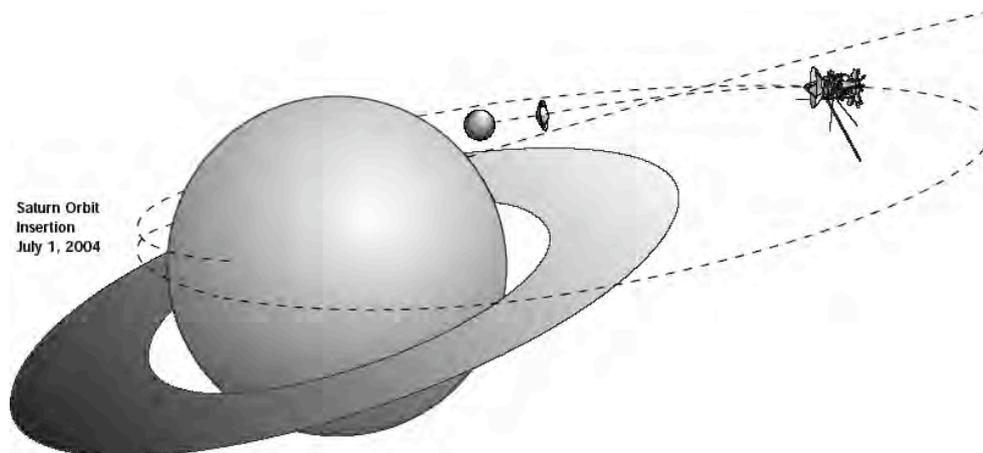
(4) Titan

(5) Icy satellites

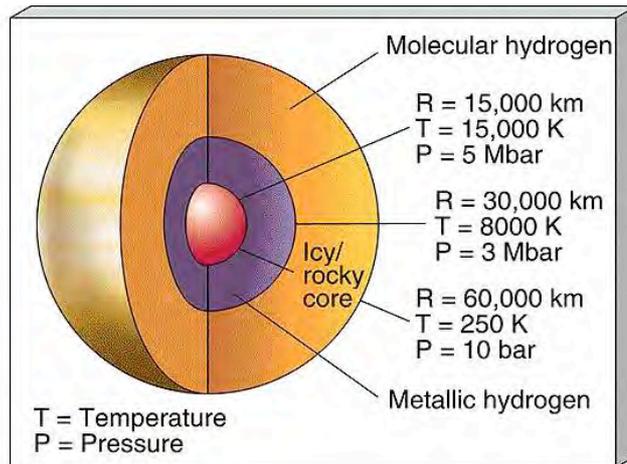
(6) Magnetosphere

• Density (0.7), Gravity & mass distribution ($J_2...$)

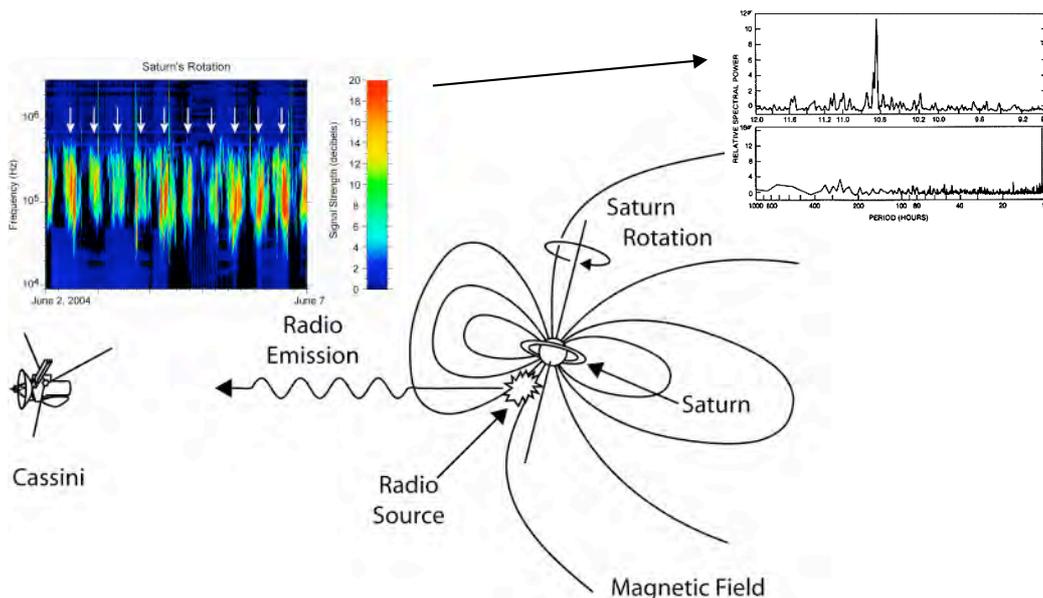
[S/C trajectories]



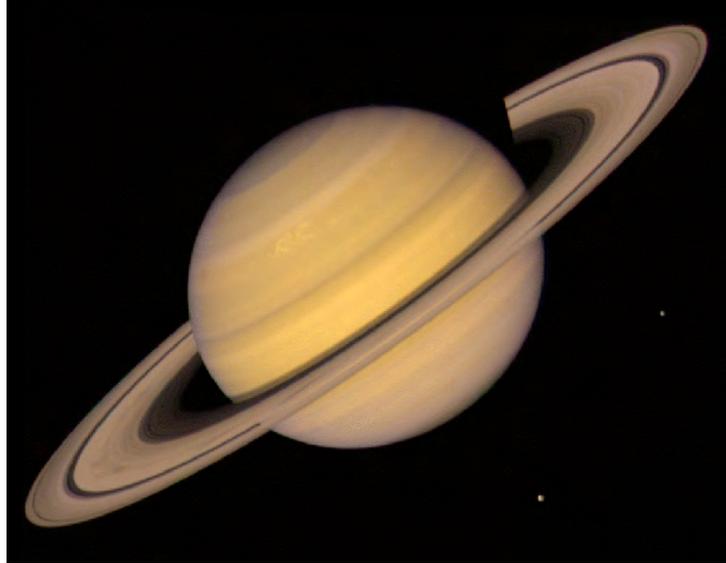
- Composition : ~94% H, 6% He [Voyager-UV Spectroscopy]
- Gaseous / liquid / metallic H,He (hot and high pressure) [Theory]
- Transition liquid / gas supercritical → no surface → $z_0 = 1$ bar level
- Emitted energy ~2 x received from Sun
- primordial condensation terminated (\neq Jupiter)
- He rain/condensation ? (6% He compared to 13% at Jupiter)



- Magnetic field ~axial → dynamo theory ? Source in metallic H ? [MAG]
- Magnetic field rotation period = 10h39.4 min [Voyager-PRA]
(from magnetospheric radio emissions modulation → reference for wind speeds)



- Radio rotation → polar flattening 9% (of fluid ball)



SOME QUESTIONS FOR CASSINI

- Internal structure *[Trajectory along many orbits]*
- Compare internal / external energy sources *[Temperature maps]*
- Magnetic anomalies ? *[MAG + Radio emissions]*

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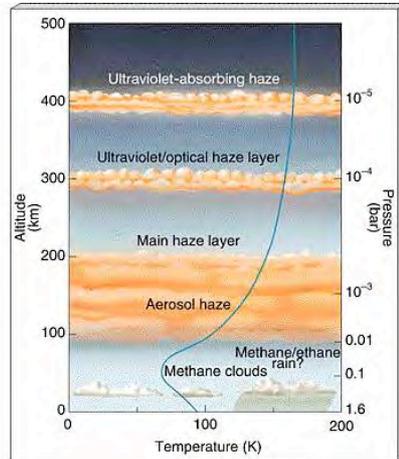
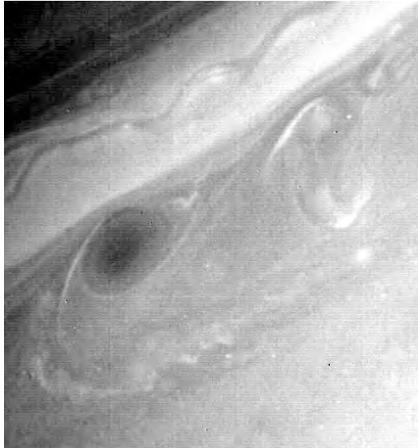
(6) Magnetosphere

- H, He, minor constituents (10^{-6}) → chemistry, colors, rain ?

[Spectroscopy]

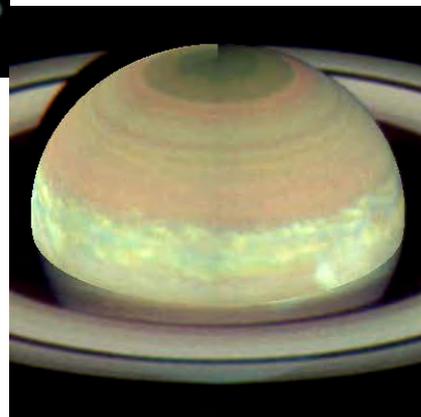


- Clouds & bands, cyclonic storms → convection + rotation
[Voyager ISS + P11 photopolarimeter]
- T(z) profile, minimum ~1 bar level
[IR spectroscopy of ≠ components]



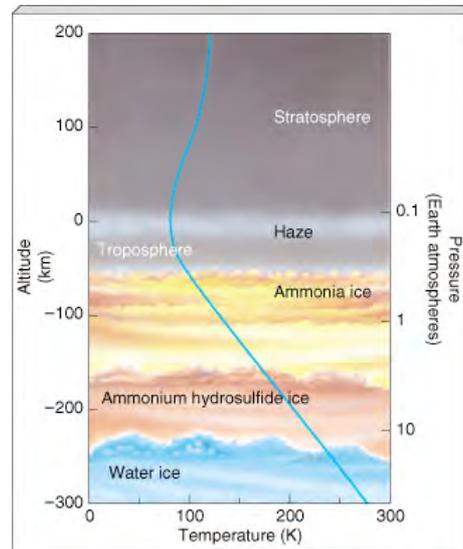
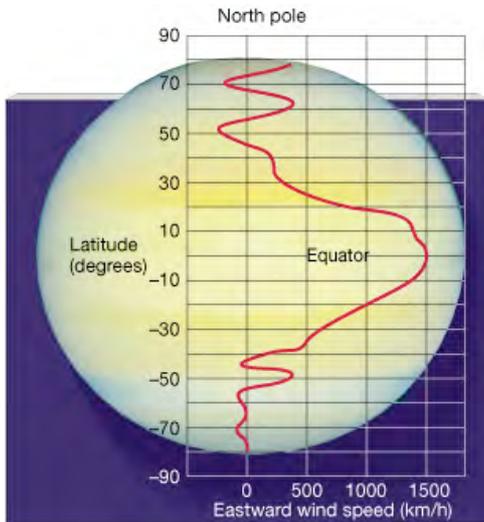
*Convection : hot gas rises, cool, form clouds, sink/rain
+ Rotation/zonal winds → generates bands, cyclones
Primary engines = ΔT (latitude, z), strong internal energy source*

- HST imaged a large storm in 1994



- Rotation period of clouds = 10h15 (equator) to 10h38 (poles)
 → winds up to 1800 km/h (= 5 × Jupiter)

[Imaging]

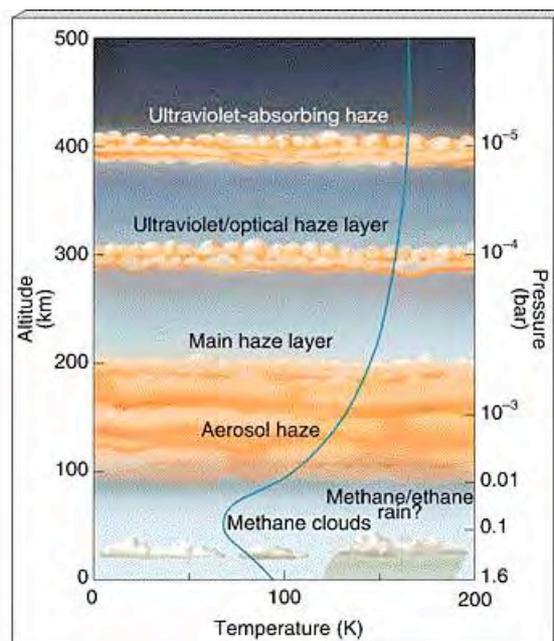


- Clouds of NH_3 , NH_4SH , H_2O ?

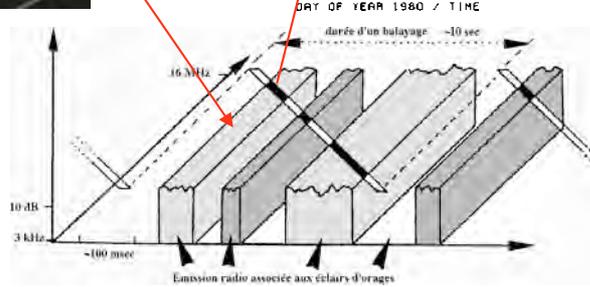
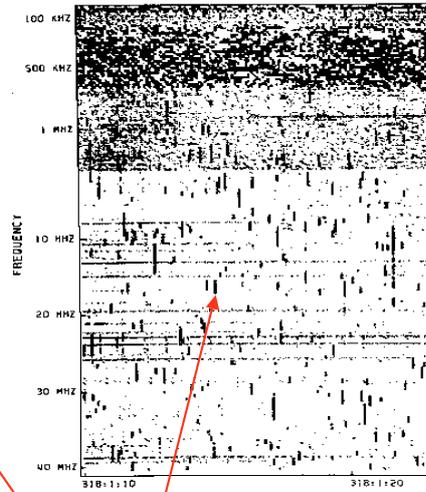
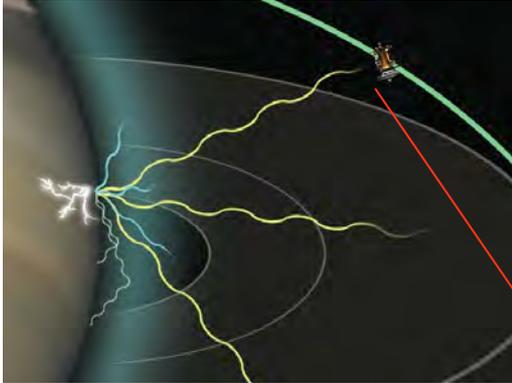
[Voyager RSS + Theory]

- Haze between 0.1 & 0.01 bar (cold planet) → dim colors

[Imaging, Spectroscopy, RSS]

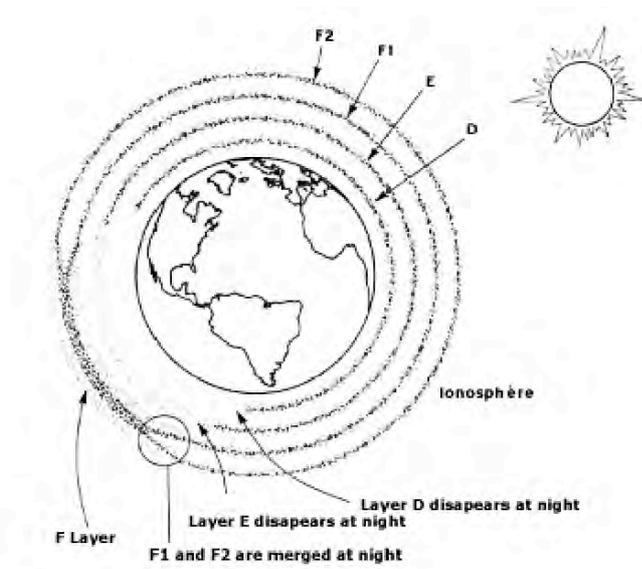


- Strong Lightning [Radio Emissions]



- Ionosphere density and temperature
- Hot thin exosphere up to 20 R_S, T=400-800 K

[Lightning + Radio Science]



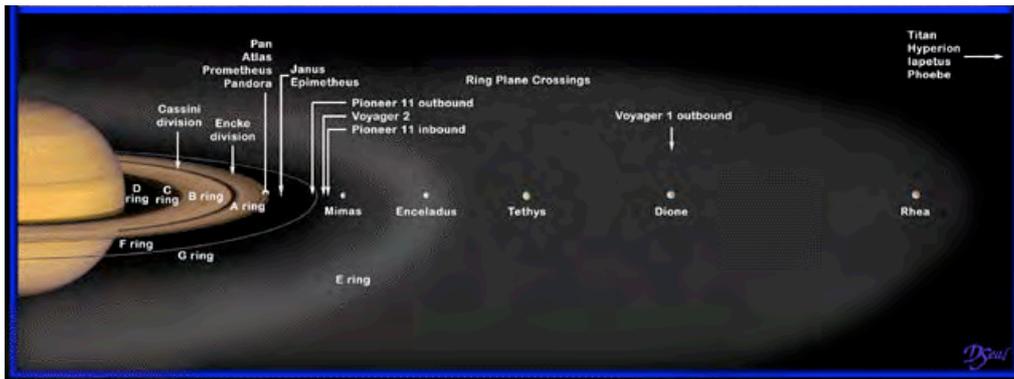
SOME QUESTIONS FOR CASSINI

- Maps of atmospheric $T(L,\lambda,z)$, clouds, composition, winds
[3 spectrographs : CIRS (molecules), VIMS, UVIS (atoms) + imaging ISS]
- Seasonal variations (white spot = NH_3 crystals ?...) *[ISS]*
- Opacity of haze & clouds *[UVIS, VIMS, RSS]*
(absorption/emission lines, scattering at sunrise/sunset)
- Map lightning *[RPWS k vector]*
- Measure He/H & D/H ratios (\rightarrow solar system origin) *[UVIS, RSS]*
- Ne, T, P, winds in ionosphere, diurnal variations *[RPWS, RSS]*

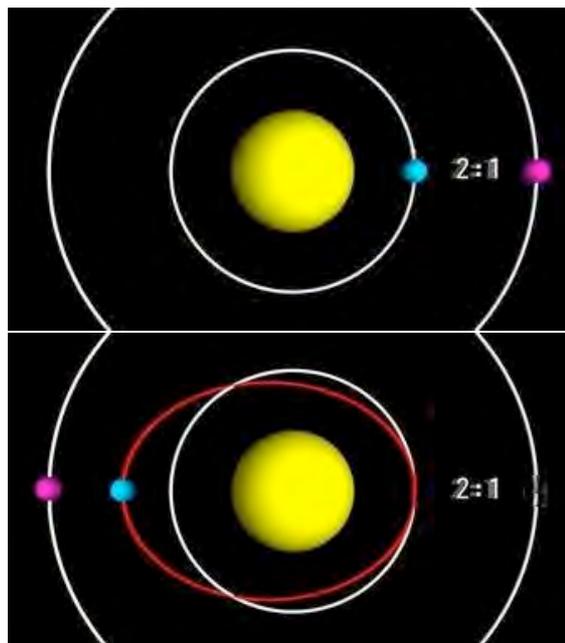
SUMMARY OF PRE-CASSINI KNOWLEDGE ABOUT SATURN & QUESTIONS TO BE ADDRESSED BY CASSINI

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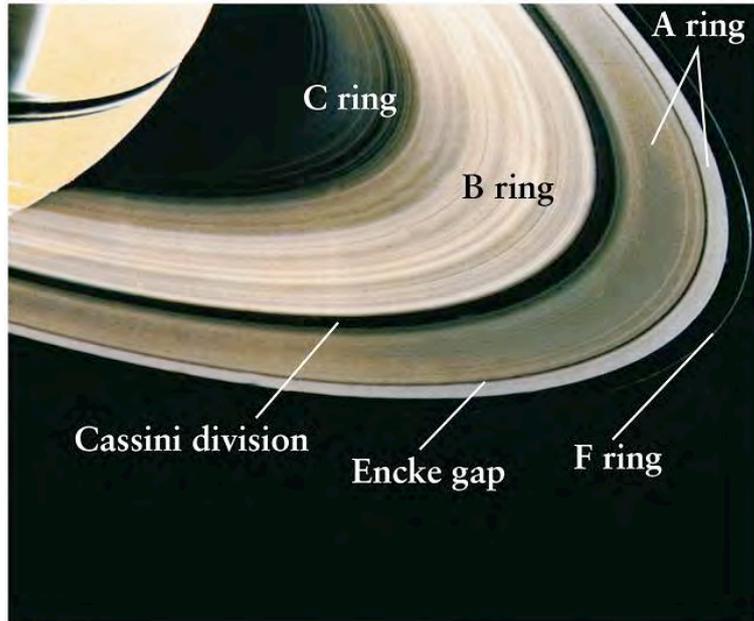
- 1866 Kirkwood finds Cassini gap resonant with Mimas also true for Keeler/Encke gap



Resonances $n:m \rightarrow$ cumulative gravitational influence



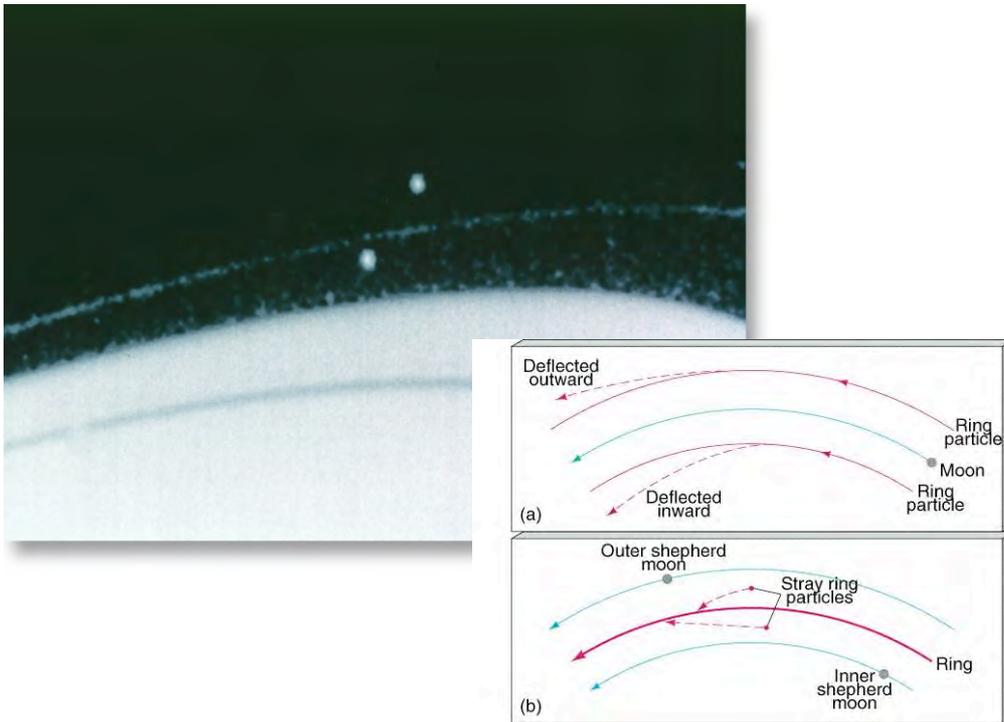
- Occultation of Iapetus by C ring → semi-transparent
- Cassini division not empty (Barnard, 1889-1908)



- 1979 Pioneer 11 discovers thin F ring [Photopolarimeter]
→ seen braided & kinked by Voyager



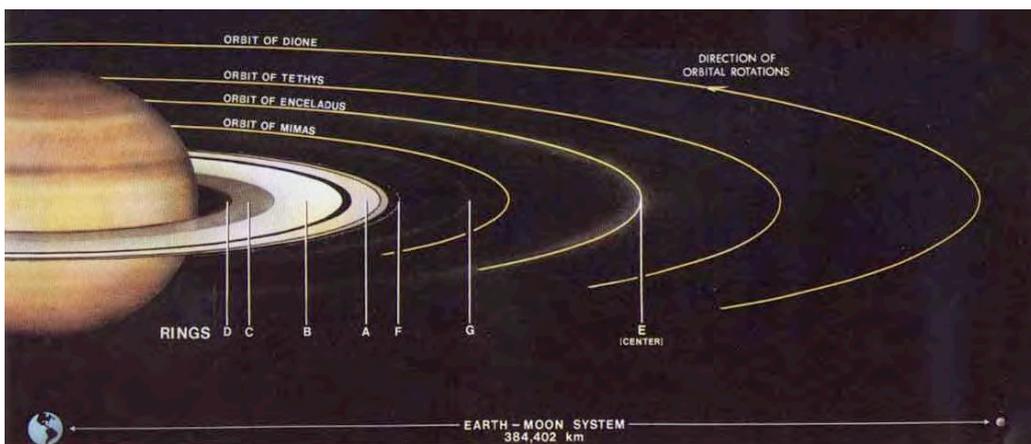
- Voyager discovered shepherds satellites



- 1980

Voyager 1 discovers D & G rings

[ISS, PPS]



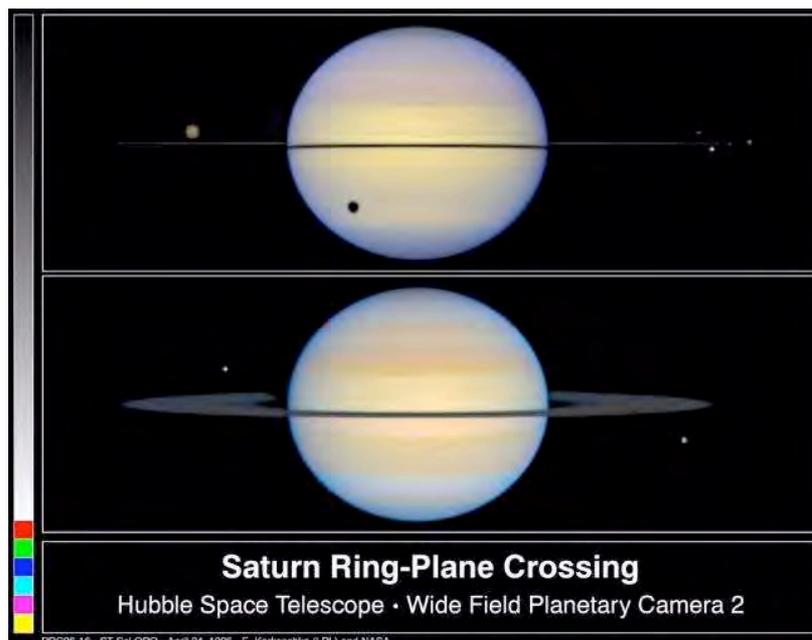
THE RINGS OF SATURN

| Ring | Distance, kilometers* | Width, kilometers |
|------|-----------------------|-------------------|
| D | 66,970 | 7,500 |
| C | 74,500 | 17,500 |
| B | 92,000 | 25,400 |
| A | 122,170 | 14,610 |
| F | 140,180 | 50 |
| G | 170,180 | 8,000 |
| E | 180,000 | 300,000 |

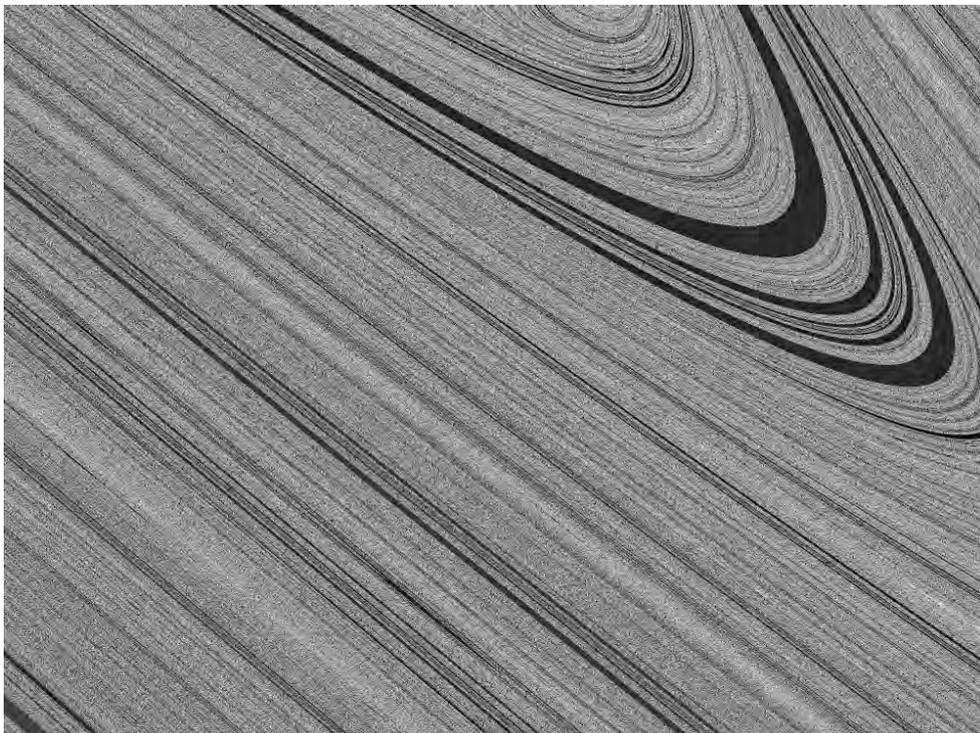
* Distance from Saturn to closest edge of ring.

- Thickness of rings ~100 m

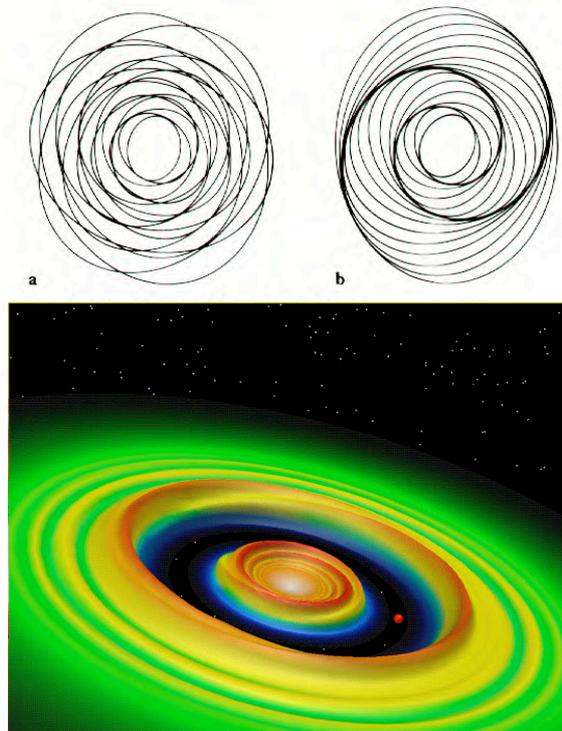
[ISS, RSS]



- 500-1000 ringlets → resonances ?
[Star occultation observed by Voyager 2, PPS, UVS]
- Eccentric & azimuthally asymmetric ringlets

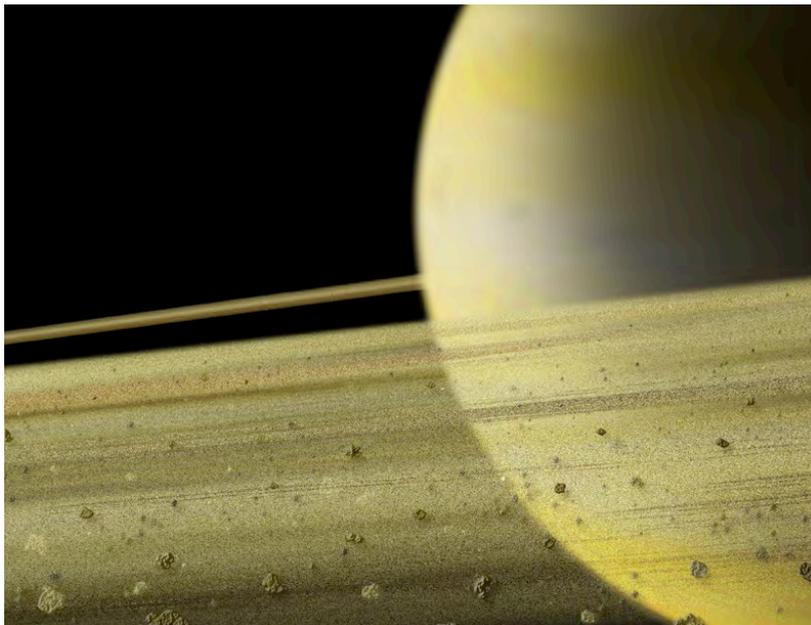


- Density waves (~spiral) → resonances

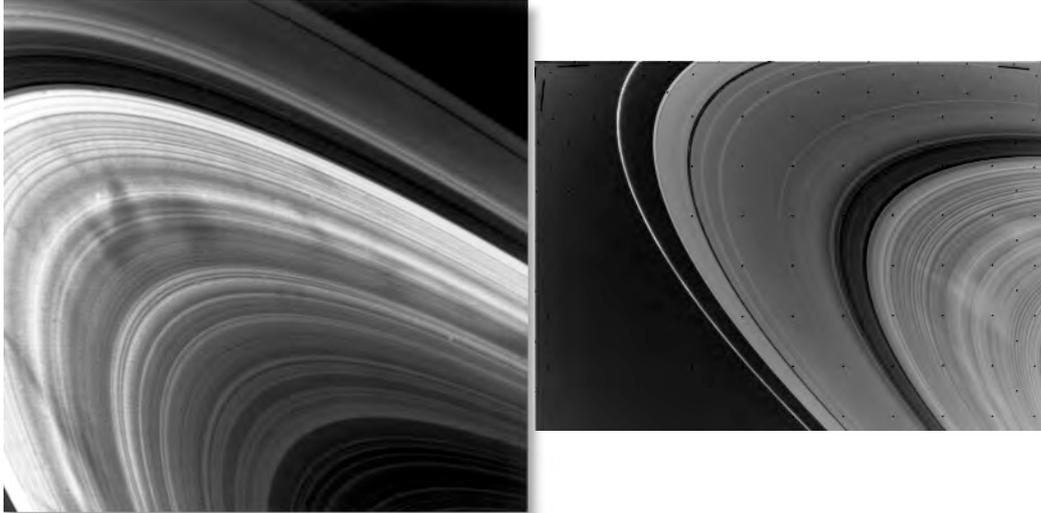


- Particle size distribution (most of the mass in cm-m range)
& measure of collision-viscosity

[RSS]



- Spokes (~radial, dark/bright features in back/forward scattered light, rapid apparition) → micron sized particles in e.s. levitation ? [ISS, PPS]

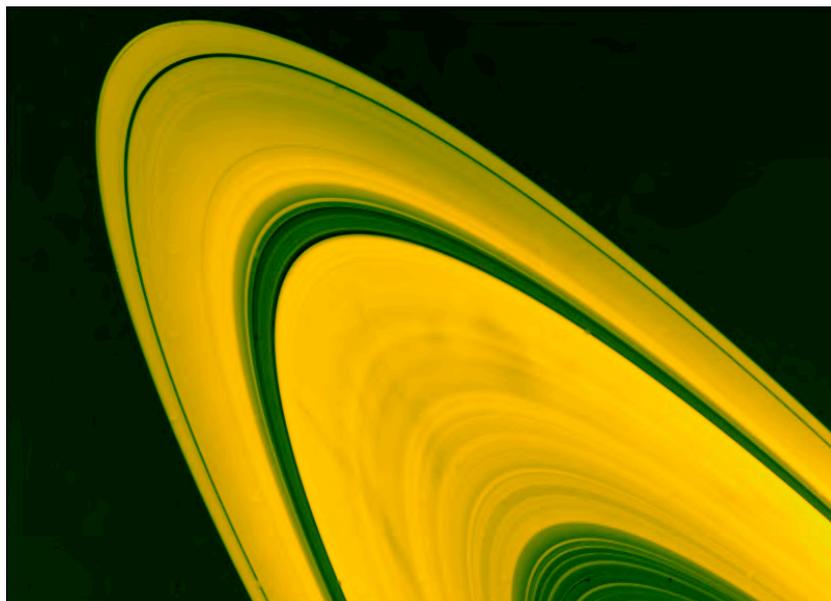


Scattering probability of a photon λ (fraction of incident light deviated) by a particle of size r :

$$P(\lambda) \propto 1/\lambda^2 \quad (r \sim \lambda) \text{ mostly forward}$$

$$P(\lambda) \propto 1/\lambda^4 \quad (r \ll \lambda) \sim \text{isotropic}$$

- Colors → composition ~ pure H₂O ice + impurities
- Differences between rings [ISS, Spectroscopy]



SOME QUESTIONS FOR CASSINI

- Why several morphologically distinct rings ?
- Complex structure (more than resonances can explain)
 - many embedded moonlets to discover ?
 - extensive monitoring for moonlets and/or density waves at ring edges
- Formation : satellite broken / not formed, or comets captured ?
- Stability ? theory predicts 40-400 Myears lifetime
 - short-lived, with recurrent formation/dissipation ?
- Why almost pure ice in spite of particle bombardment ?
- Spoke formation ? Relation to Magnetosphere & Radio emission ?
- Relation between ring composition and satellite composition (e.g. is Enceladus the source of E-ring ?)

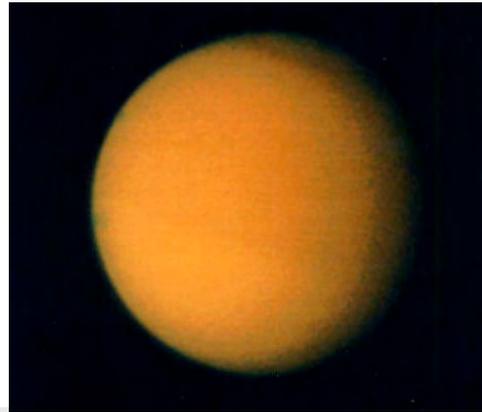
[INMS, CDA, CAPS measurements in many geometries, illuminations, phase angles]

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- (1) Saturn's interior**
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- Orbit at 20.3 R_S from Saturn

→ *Voyager 1 flyby at 4000 km*

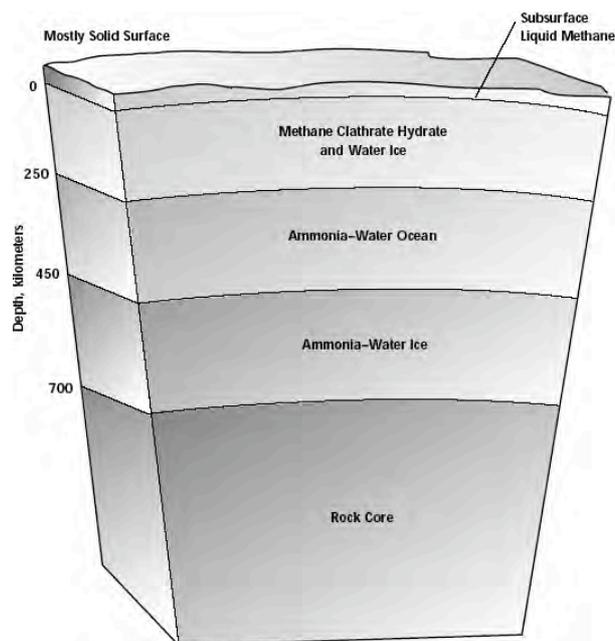


SIX GIANT SATELLITES

| Satellite (Planet) | Titan (Saturn) | Moon (Earth) | Io (Jupiter) | Europa (Jupiter) | Ganymede (Jupiter) | Callisto (Jupiter) |
|---|----------------|--------------|--------------|------------------|--------------------|--------------------|
| Distance from Parent, kilometers | 1,221,850 | 384,400 | 421,600 | 670,900 | 1,070,000 | 1,883,000 |
| Rotation Period, days | 15.945 | 27.322 | 1.769 | 3.551 | 7.155 | 16.689 |
| Radius, kilometers | 2575 | 1738 | 1815 | 1569 | 2631 | 2400 |
| Average Density, grams per cubic centimeter | 1.88 | 3.34 | 3.57 | 2.97 | 1.94 | 1.86 |

- Density 1.88 → 50% rock, 50% H₂O ice

[S/C trajectory]



- Atmosphere : N_2 + a few % CH_4

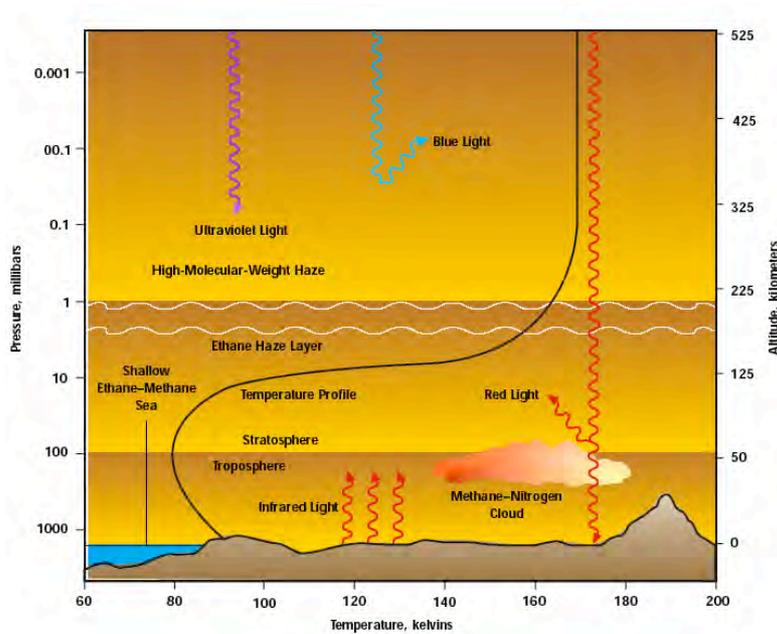
[Spectroscopy]

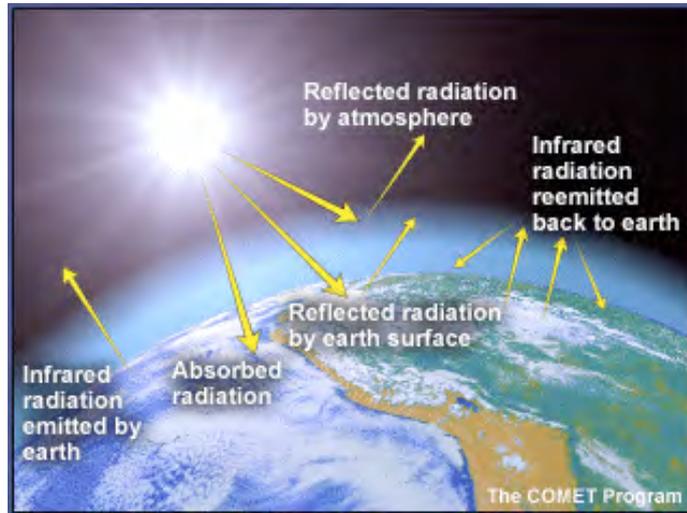
CONSTITUENTS OF THE TITAN ATMOSPHERE

| Chemical Constituent | Common Name | Atmospheric Concentration |
|----------------------|------------------|---------------------------|
| N_2 | Nitrogen | 90–97 percent |
| Hydrocarbons | | |
| CH_4 | Methane | 2–10 percent |
| C_2H_2 | Acetylene | 2.2 parts per million |
| C_2H_4 | Ethylene | 0.1 parts per million |
| C_2H_6 | Ethane | 13 parts per million |
| C_3H_8 | Propane | 0.7 parts per million |
| Nitriles | | |
| HCN | Hydrogen cyanide | 160 parts per billion |
| HC_3N | Cyanoacetylene | 1.5 parts per billion |

- Thick atmosphere ($P_{\text{surface}} = 1.5 \text{ bar}$) → Greenhouse effect, $T_{\text{surface}} = 94 \text{ K}$

[Star occultation, 1989]





Greenhouse effect :

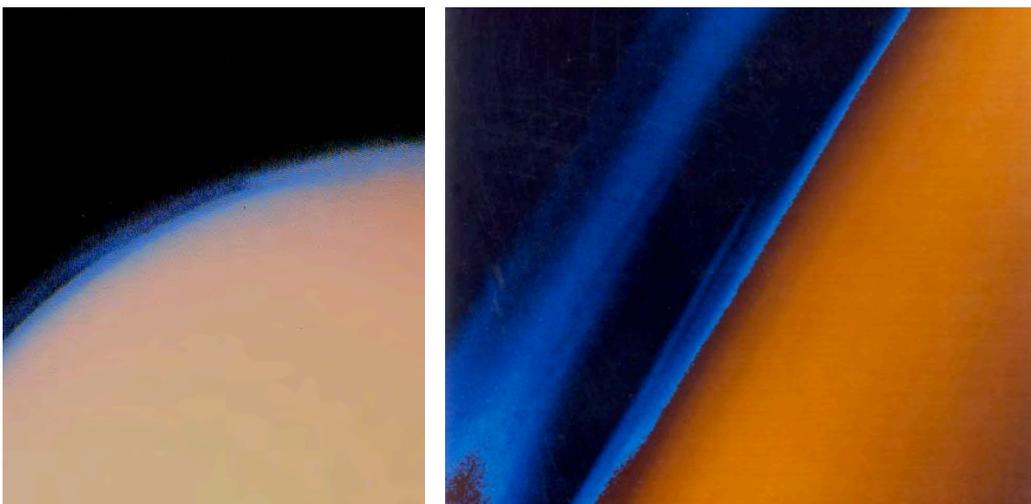
Thick atmosphere lets Visible light In, prevents IR to exit

→ heats surface → reradiates

→ T_{max} at low & high z, T_{min} at mid z

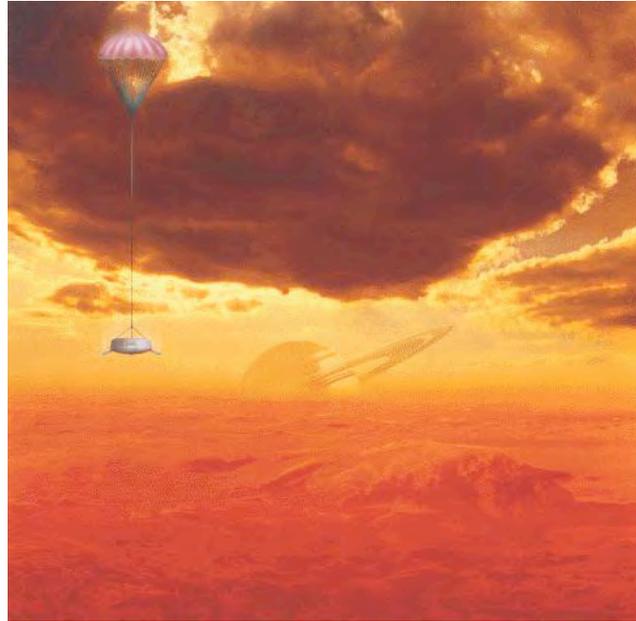
- Photochemistry (+ bombardment by magnetospheric electrons)
 → C_nH_n smog & haze (500 km > clouds)
 → condense/polymerize + rain ? → organic compounds on soil ?

[Imaging, Spectroscopy + Theory, Modelisation, Lab. Experiments]



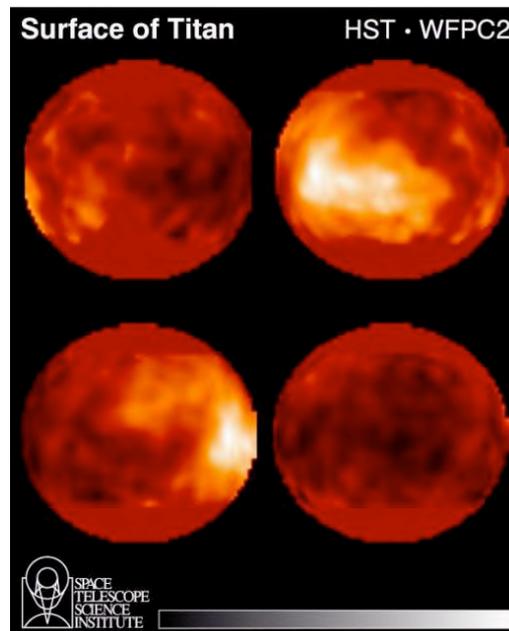
Smoggiest world in solar system, 1/1000 solar luminosity / Earth !

- Solid surface ? Ocean (up to 1 km deep) ? Lakes or subsurface (in sputtered regolith) reservoirs of CH_4 , C_2H_6 ... ?



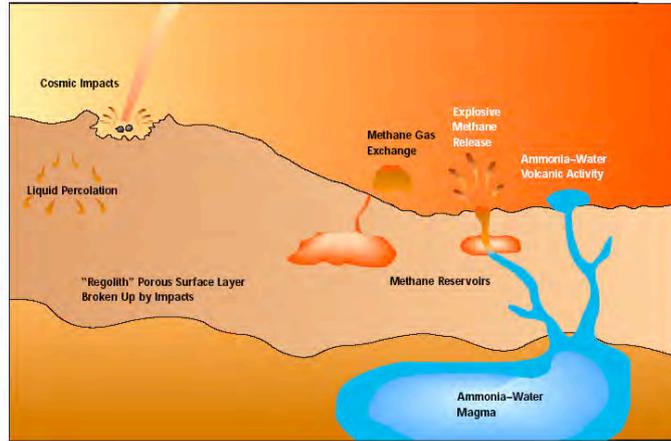
- Ground-based observations → no global ocean, bright spot

[Earth based RADAR + NIR/AO + HST/WFPC2]



- Ice tectonics ? $H_2O + NH_3$ magma, possibly erupting ?

*15% more NH_3 at 10 AU than at 5 AU ; NH_3 ~antifreeze
 → liquid below H_2O ice surface → favours cold tectonics ?*



- ΔT equator-poles = 15K → winds, jet streams at high altitude ?

[IR spectroscopy]

- No lightning ?

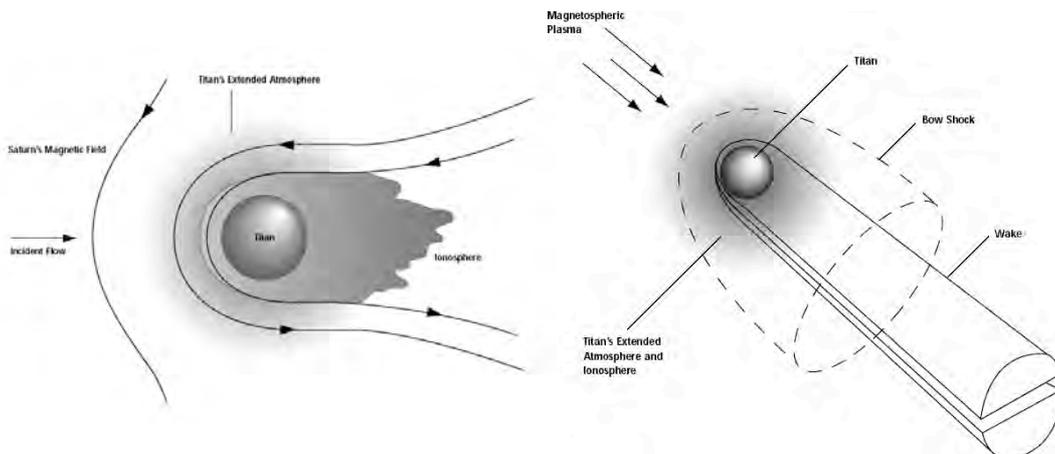
[Voyager-PRA]

- No magnetic field ? ($B_{\text{equator}} < 4 \text{ nT}$)

[MAG]

- Plasma wake (interaction with magnetosphere)

[Voyager-PLS, PWS]



SOME QUESTIONS FOR CASSINI

- Structure of interior ? differentiated ? *[S/C path along 40 close flybys]*
- Origin / evolution of atmosphere ? *[Spectroscopy UVIS, VIMS, CIRS]*
- N₂ primordial or from cometary NH₃ ? *[Ar/N₂ concentrations]*

Ar & N₂ have same condensation Temperature → if N₂ primordial, Ar/N₂ ratio should have the solar nebula value ; Voyager suggests 6× less (but poor sensitivity)

- H escapes in C_nH_n reactions → irreversible → lost in 10⁷ years
→ source of C_nH_n cycle ? ocean ?
- Clouds, rain ? (Voyager → supersaturated atmosphere) *[RSS]*
- Lightning ? *[RPWS, ISS, HP instruments]*

SOME QUESTIONS FOR CASSINI (continued)

- Nature of surface ? Ocean, continent size plateau ?
[long (IR) wavelength imaging - VIMS]
- Ice tectonics, Craters, Erosion → geological activity ?
- Prebiotic chemistry ? Titan = Primitive Earth in fridge ?
(but no liquid water except perhaps temporarily after large impacts)
[Spectroscopy, HP instruments]
- Interaction with dust & ring particles ? *[CDA]*
- B field ? Interaction with Magnetosphere (transsonic) or Solar Wind
(supersonic) at various locations *[MAG, INMS, RPWS, UVIS]*
- Source of Magnetospheric plasma ? Nitrogen torus ? *[CAPS, INMS]*
- High energy particle environment (bombardment) *[MIMI]*

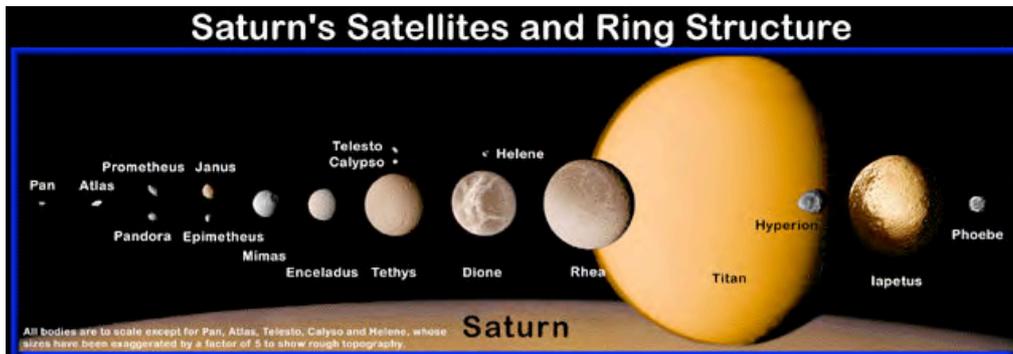
Huygens Probe instruments :

- HASI → atmospheric properties (T, density, P, electrical)
- Radar Altimeter → surface properties (Δz , roughness)
- GCMS → chemical comp. analysis
- ACP → captures/vaporizes aerosols → sends to GCMS
- DISR → opacity, clouds, surface features (solar energy deposited at each altitude, reflectance)
- Doppler Wind Experiment → trajectory, winds, turbulence, liquid waves (+VLBI)
- SSP → surface solid/liquid properties



SUMMARY OF PRE-CASSINI KNOWLEDGE ABOUT SATURN & QUESTIONS TO BE ADDRESSED BY CASSINI

- (1) Saturn's interior
- (2) Saturn's atmosphere
- (3) Rings
- (4) Titan
- (5) Icy satellites
- (6) Magnetosphere

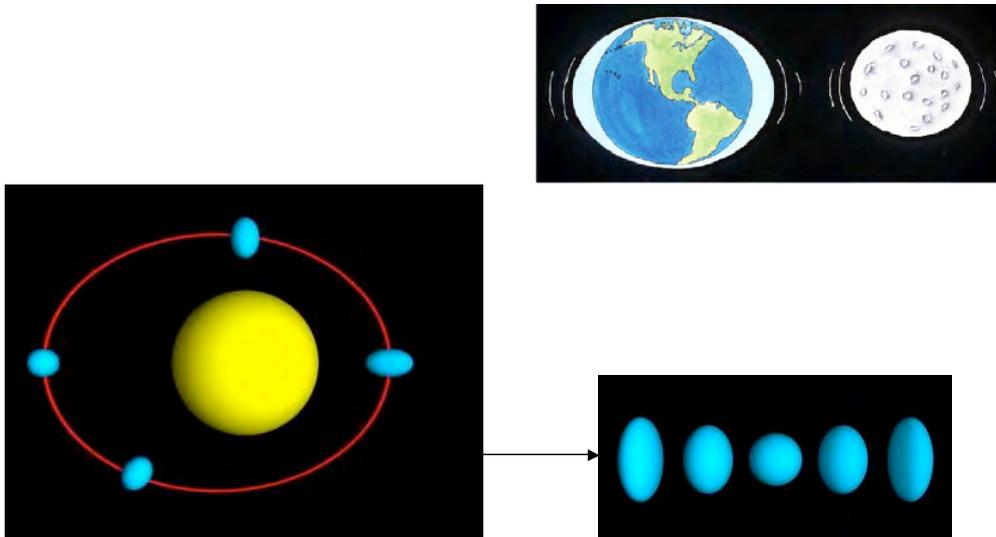


- After Voyager 17 satellites known (+1 found in 1990 using 1981 data)
[Voyager trajectories, Imaging, Crater counts + Spectroscopy]

THE KNOWN MOONS OF SATURN

| Moon | Diameter, kilometers | Distance, kilometers | Orbital Period, days | Year Discovered: Discoverer |
|------------|----------------------|----------------------|----------------------|-----------------------------|
| Pan | 20 | 133,580 | 0.56 | 1990: Showalter |
| Atlas | 30 | 137,670 | 0.60 | 1980: Terrile |
| Prometheus | 100 | 139,350 | 0.61 | 1980: Collins |
| Pandora | 90 | 141,700 | 0.63 | 1980: Collins |
| Epimetheus | 120 | 151,450 | 0.69 | 1966: Walker |
| Janus | 190 | 151,450 | 0.69 | 1966: Dolfus |
| Mimas | 392 | 185,520 | 0.94 | 1789: Herschel |
| Enceladus | 500 | 238,020 | 1.37 | 1789: Herschel |
| Tethys | 1060 | 294,660 | 1.89 | 1684: Cassini |
| Telesto | 30 | 294,660 | 1.89 | 1980: Smith |
| Calypso | 26 | 294,660 | 1.89 | 1980: Smith |
| Dione | 1120 | 377,400 | 2.74 | 1684: Cassini |
| Helene | 32 | 377,400 | 2.74 | 1980: Laques and Lecacheux |
| Rhea | 1530 | 527,040 | 4.52 | 1672: Cassini |
| Titan | 5150 | 1,221,830 | 15.94 | 1655: Huygens |
| Hyperion | 290 | 1,481,100 | 21.28 | 1848: Bond |
| Iapetus | 1460 | 3,561,300 | 79.33 | 1671: Cassini |
| Phoebe | 220 | 12,952,000 | 550.40 | 1898: Pickering |

- All tidally locked in synchronous rotation except Hyperion & Phoebe

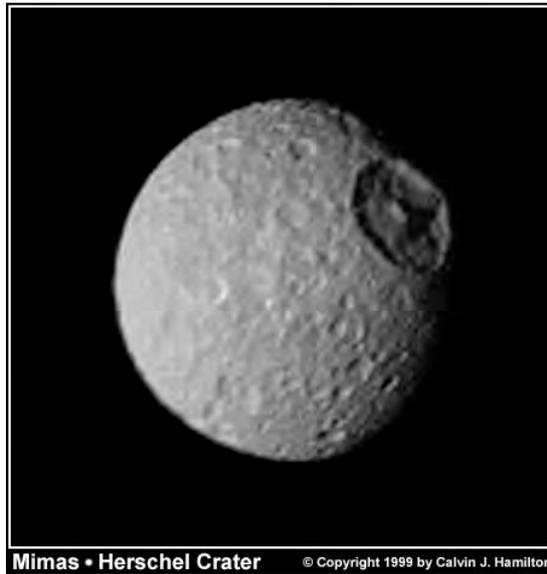


Tidal forces → energy dissipation → locking of smaller body in synchronous rotation (in a few 10^6 years)

- Most sustained geological activity during $\geq 10^9$ years (differentiation, eruptions, resurfacing)
 - Engines = primordial gravitational energy, tidal/resonant interactions, cratering, or radioactive decay
 - Presence of hydrated NH_3 , CH_4 in H_2O clathrate may increase geological activity
- Not all geologically dead : Is Enceladus still active ?
- Sputtering by micrometeorites or energetic particles
 - modify surface properties (brighter if fresh material excavated)
 - leading face of satellites often brighter (*cf. Iapetus*)
- No atmosphere (thermal escape)

- 6 medium sized satellites :

- Mimas : large impact crater (Herschel) + breaks
→ little viscous relaxation, episodes of resurfacing



- Enceladus : bright with streaks (~Ganymede), pure water ice ?
extensive resurfacing through volcanism ?



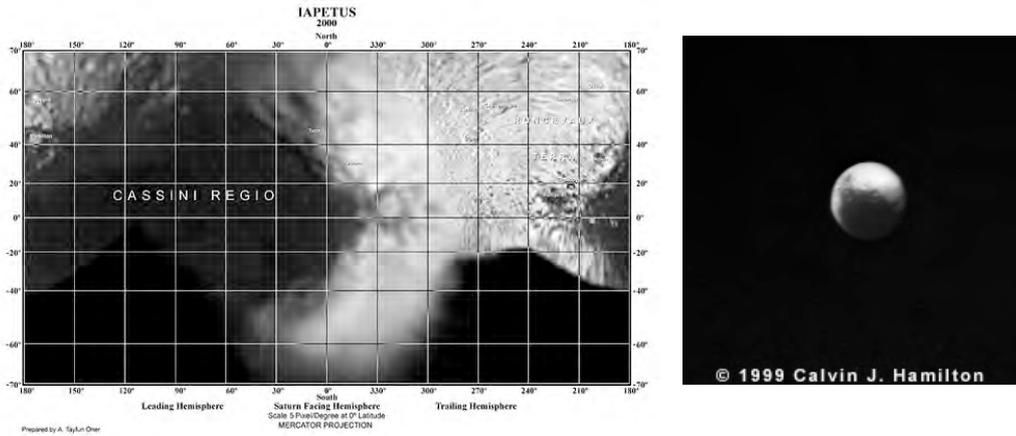
- Tethys : largest impact crater (Odysseus), huge trench (Ithaca chasma)



- Dione : more dense, diverse surface morphology
 - internal activity ? Auroral radio source modulation ?



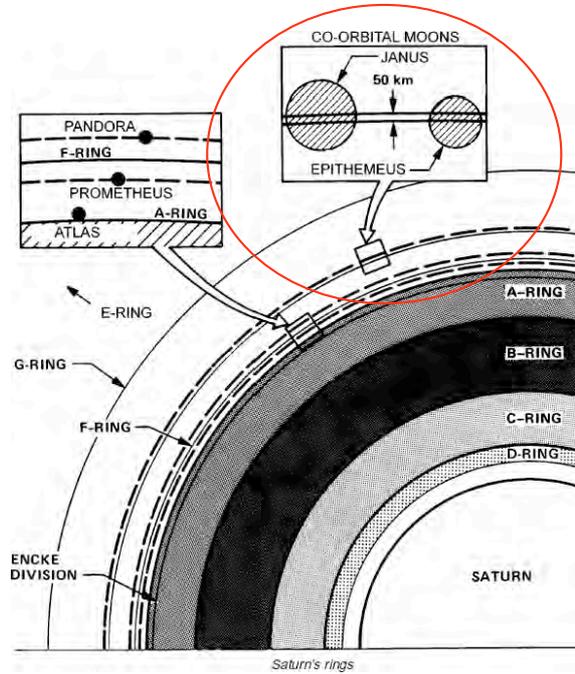
- Iapetus : bright/dark faces, highly inclined orbit



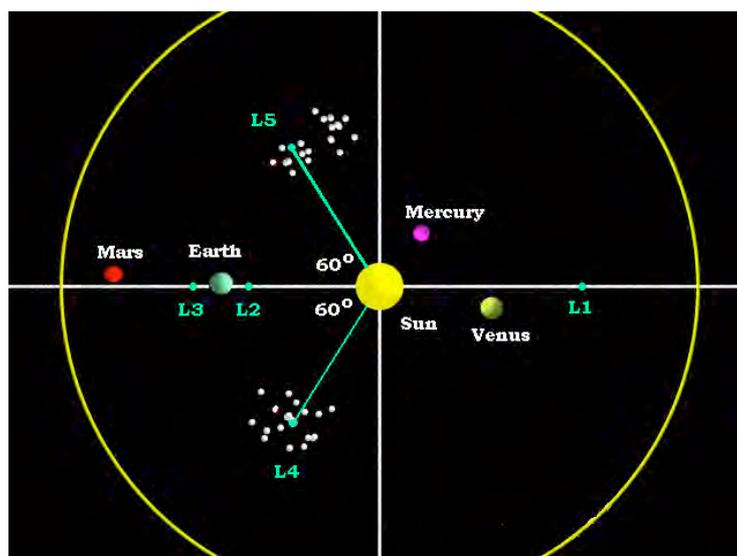
- Small satellites with irregular shape :
 - 3 shepherds : Prometheus & Pandora for F ring
Atlas at outer edge of A ring



- Co-orbital Janus & Epimetheus (discovered in 1966 & 1978) crossing with orbit exchange every 4 years



- Calypso, Helene & Telesto at Lagrangian points (60°) of larger satellites (~trojan asteroids / Earth or Jupiter)



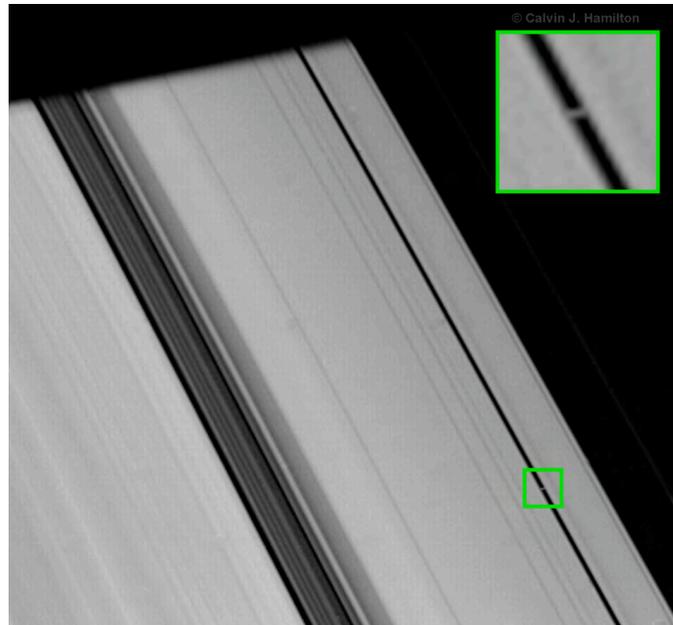
- Phoebe : outermost, dark and highly inclined retrograde orbit
 - captured asteroid ?
 - simple asynchronous rotation of 9 hours



- Hyperion : larger but more irregular than Phoebe, chaotic rotation
 - bombardment / fragmentation ?
 - perhaps recently knocked out of tidal locking



- Pan : embedded in Encke gap in A ring , last discovered



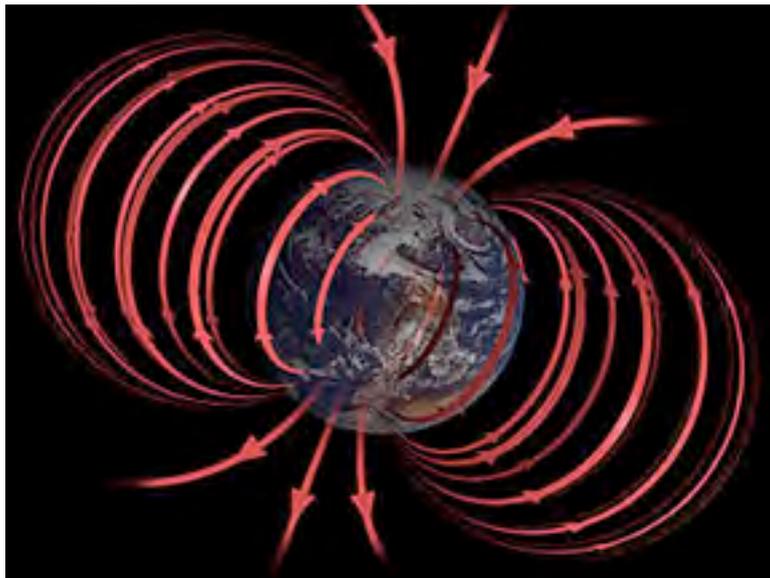
SOME QUESTIONS FOR CASSINI

- New satellites, densities, gravitational interactions with rings
[S/C path along many close flybys]
- Morphology, geology, crater counts, relation to composition
→ structure, differentiation ? *[ISS]*
- Surface compositions (volatiles, hydrates, clathrates, impurities)
[VIMS, UVIS, CIRS]
- Origin of dark material on Iapetus, Hyperion, Phoebe
→ unprocessed primordial or organic ?
- Comparison of rings, dust, satellite composition → origin ? *[CDA]*
- Enceladus = source of E ring ? still active ? (look for geysers, plumes)
- Origin of Phoebe ? Captured from Kuiper belt ?
- Relations satellite – magnetosphere (sources ?) *[CAPS, INMS, RPWS]*
- Radio emission modulation by Dione ? *[RPWS, MAG]*

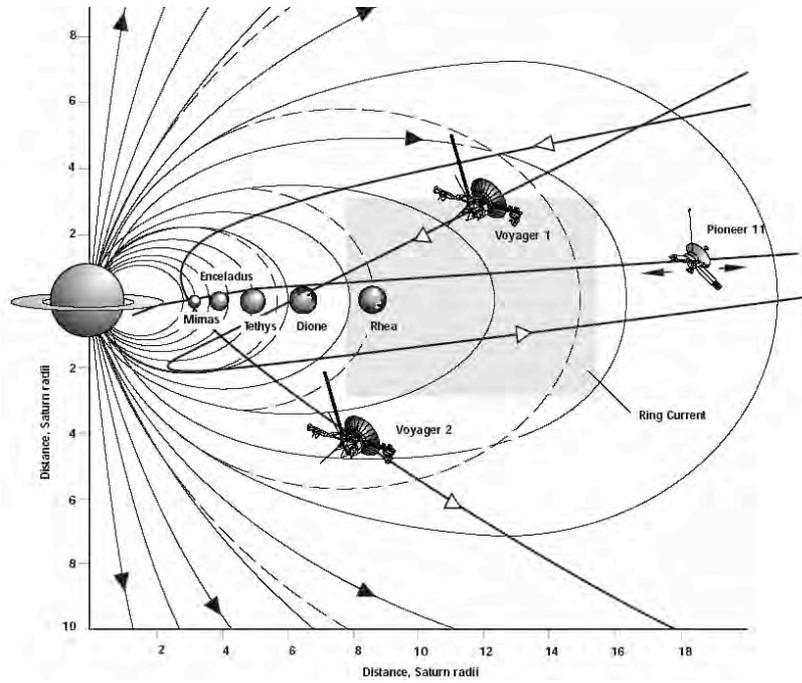
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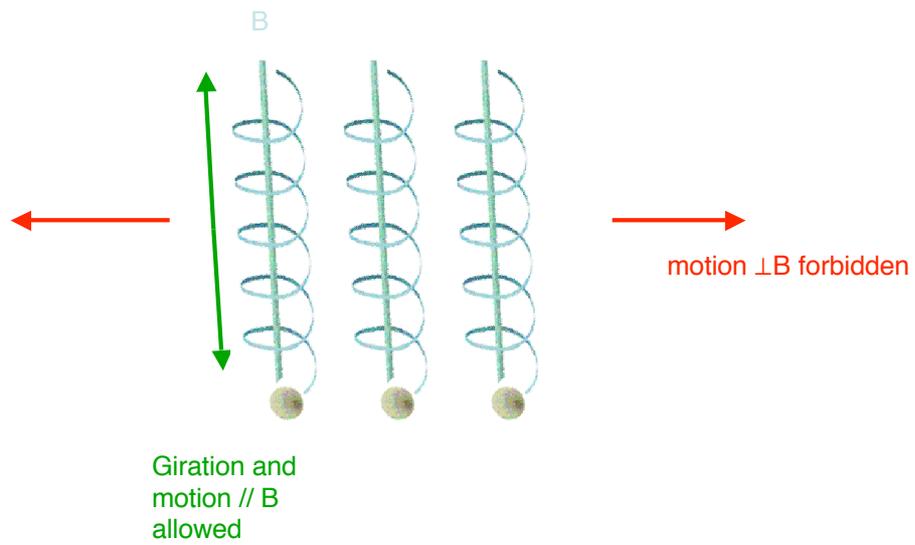
Planetary magnetic fields are dipole-like



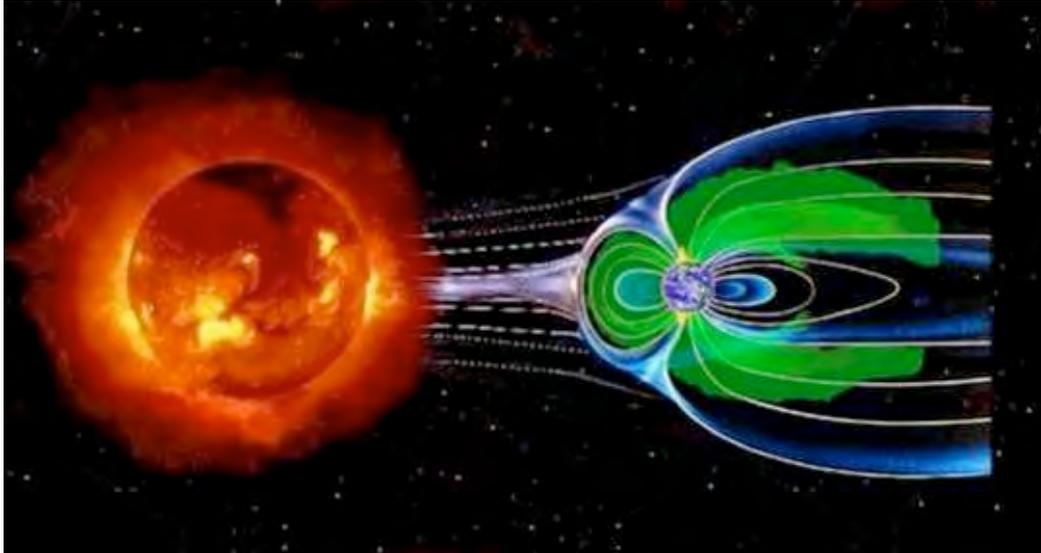
- Saturn's field with polarity opposed to Earth (same as Jupiter), ~axial



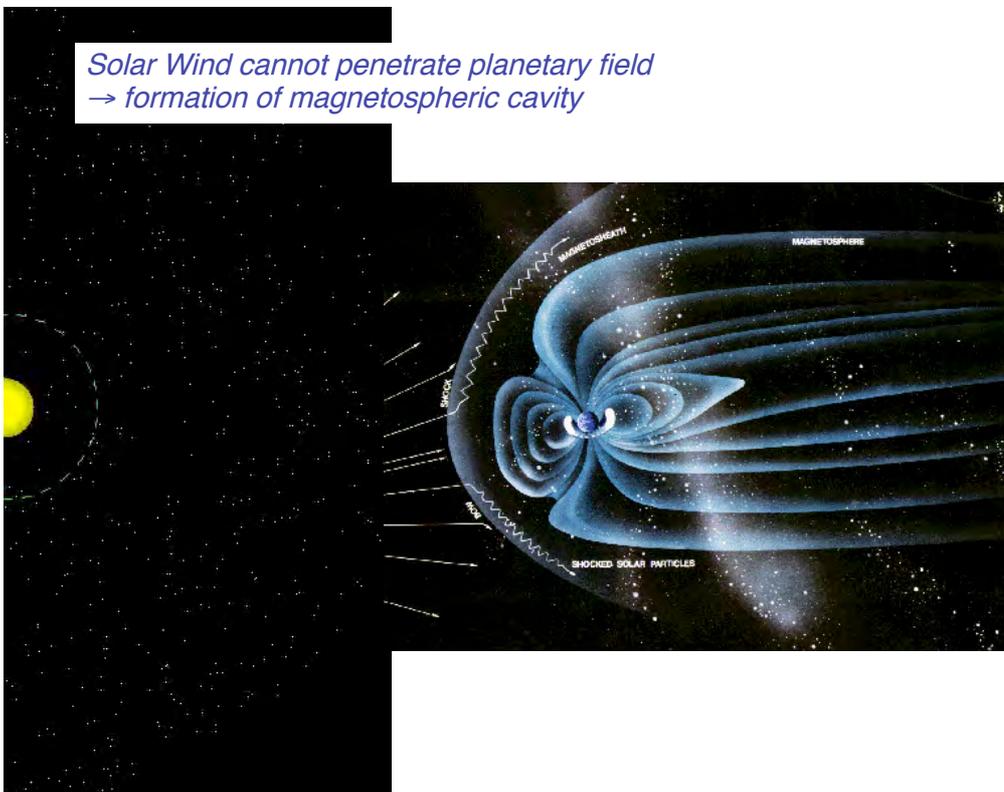
Constrains the motion of charged particles



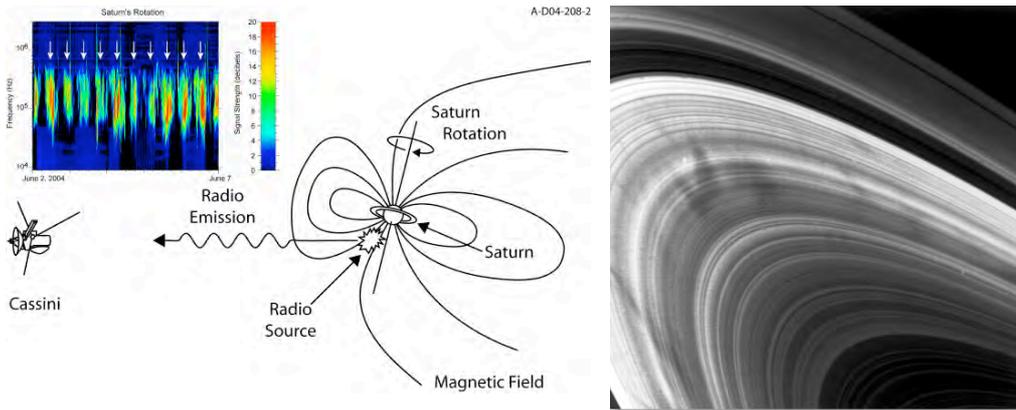
*Solar Wind cannot penetrate planetary field
→ formation of magnetospheric cavity*



*Solar Wind cannot penetrate planetary field
→ formation of magnetospheric cavity*

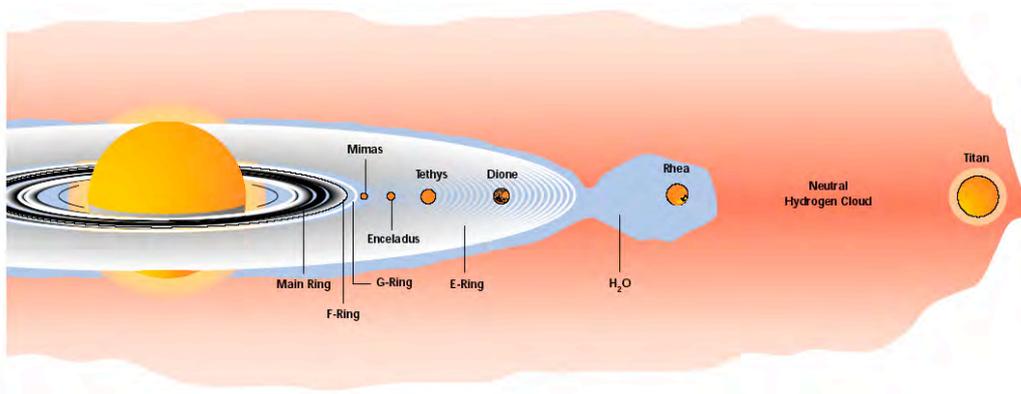


- Saturn's magnetosphere size = 20 x Earth's magnetosphere
[Waves & particles instruments + UV on Pioneer 11 and Voyager 1 & 2, + HST]
- B axial forbidden by dynamo theory
 + rotational modulation of radio emission, energetic electron spectra, spokes formation ...

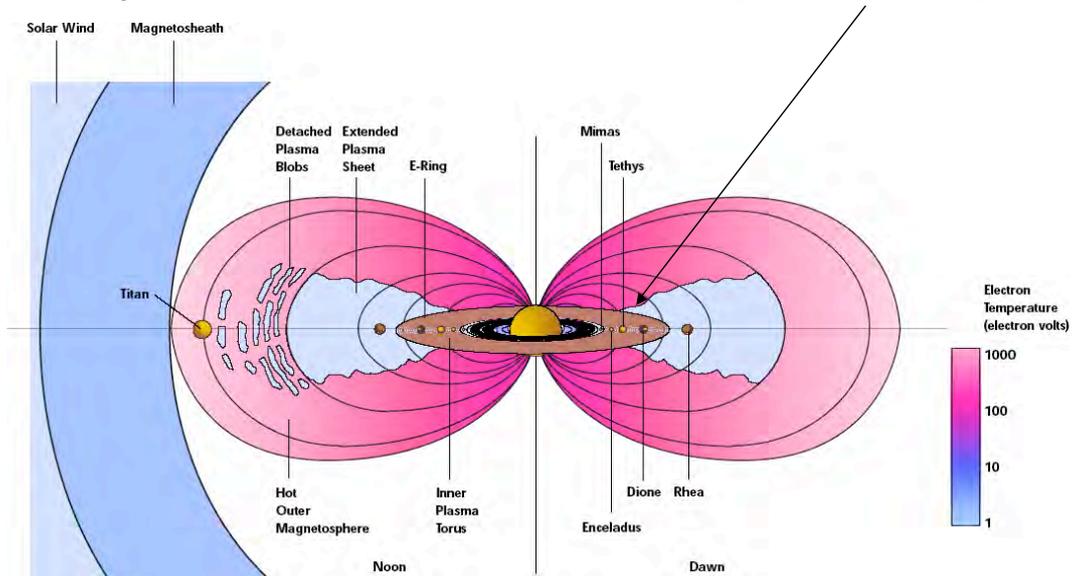


- 1% variations of radio period measured by Ulysses !

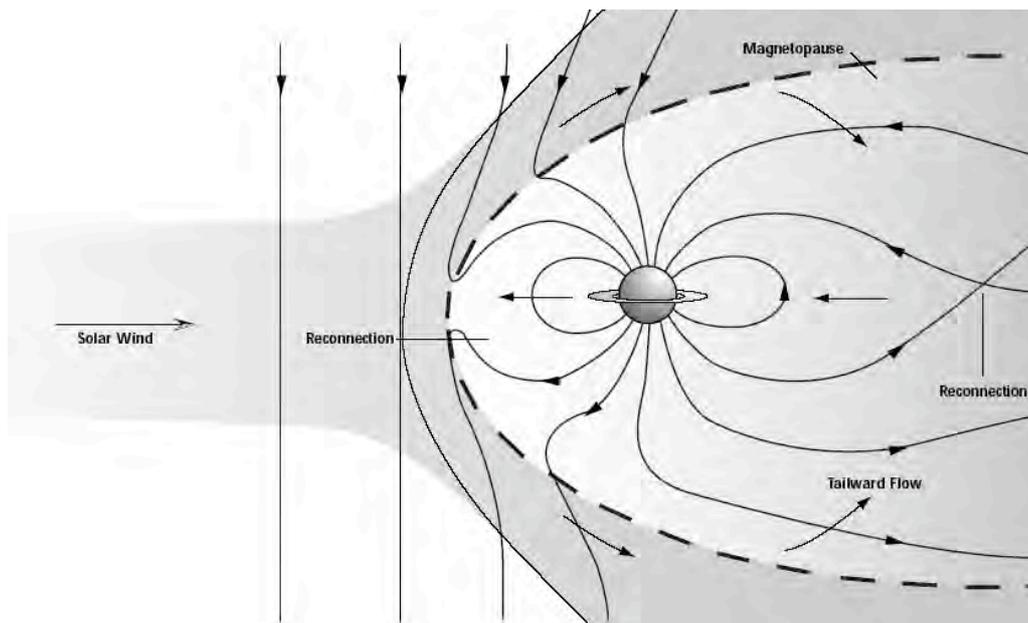
- Many sources of plasmas
 (escape, sputtering from rings and satellites, Titan, solar wind, ionosphere)
 2 at Earth, 3 at Jupiter



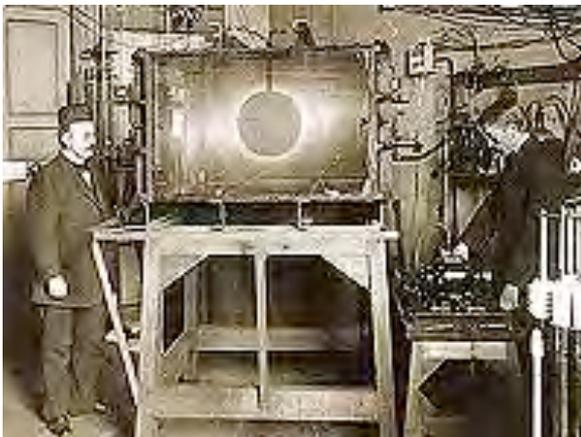
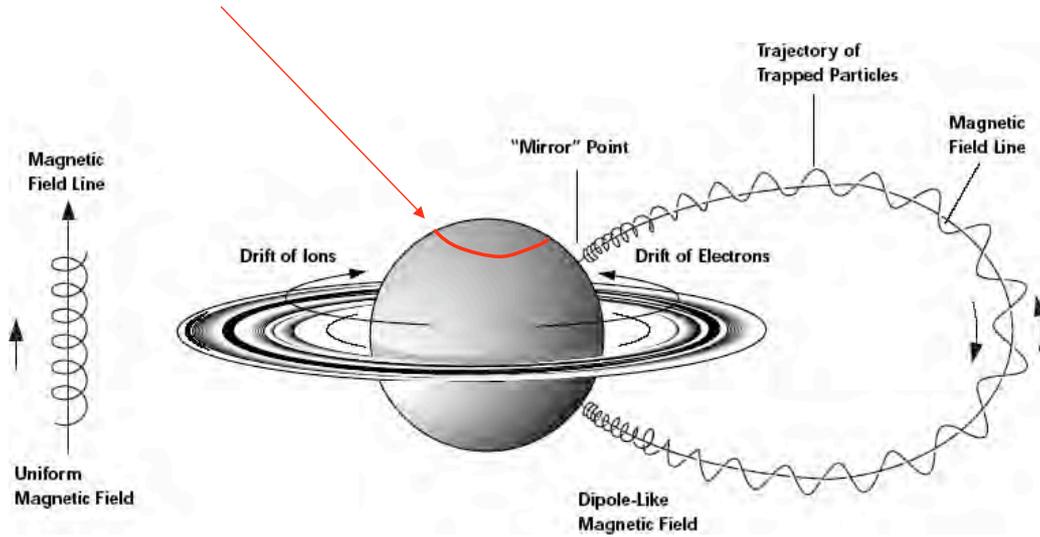
- Magnetospheric structure :
 - quiet inner MS (interaction with ring material ?), with Dione-Tethys O torus
 - hot outer variable MS (rotational energy ?)
- Rings absorption → radiation belt outwards of main rings (around E,F,G rings)



- Circulation of plasma in magnetosphere (meridian)



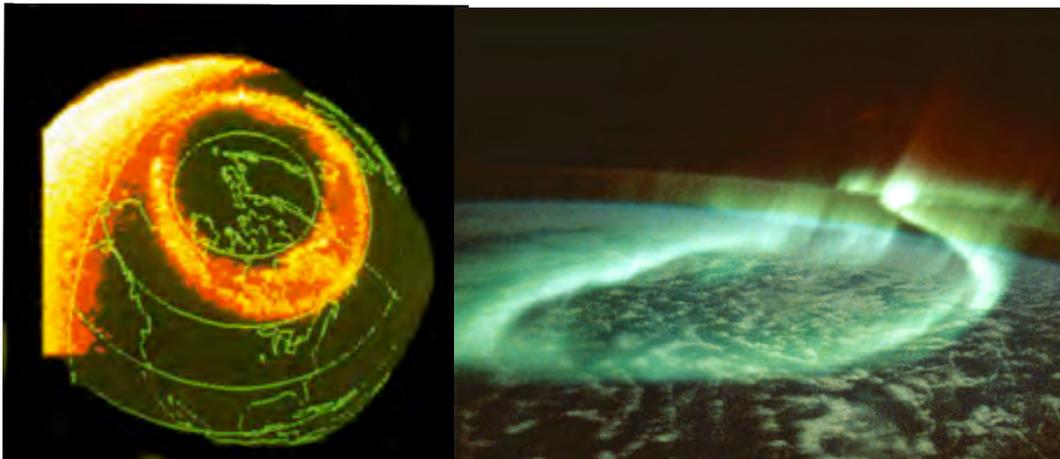
- Particle precipitation & collision at high latitudes
→ aurora (UV, IR, radio)



laboratory experiment by Kristian Birkeland (~1910) : « Terella »



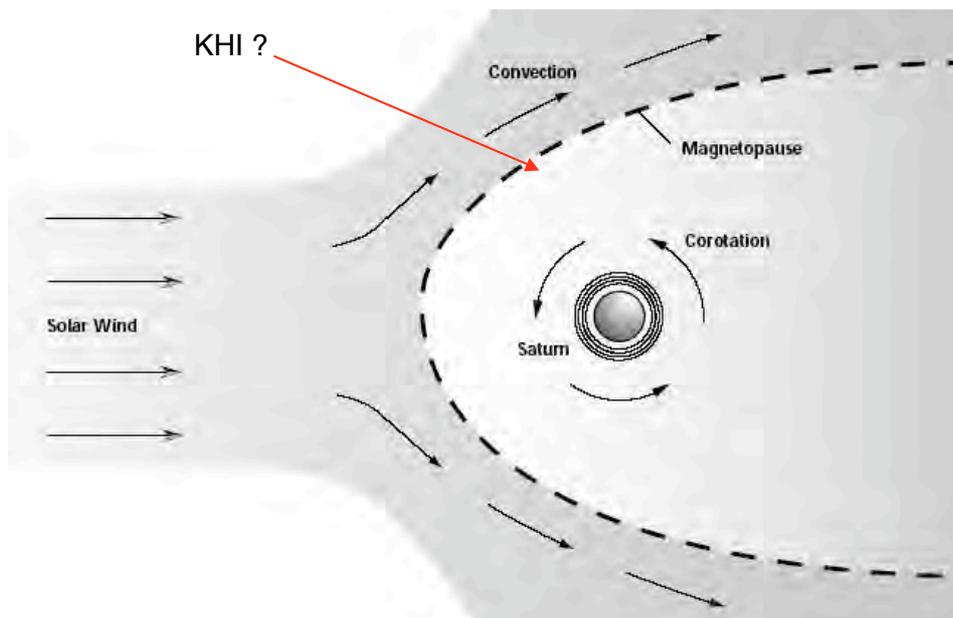
On Earth ...



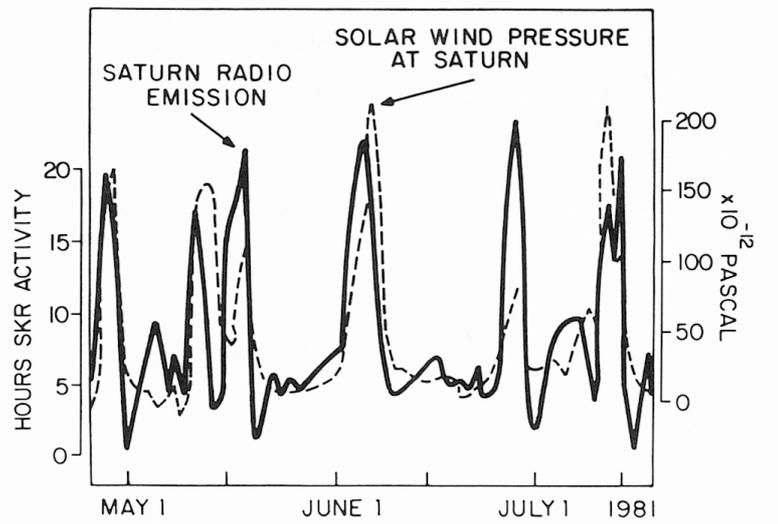
... and on Saturn



- Circulation of plasma in magnetosphere (equatorial)
→ corotation should dominate ...



... but tight control of solar wind on magnetospheric activity and aurora
(radio extinctions when Saturn in Jovian magnetotail)



SOME QUESTIONS FOR CASSINI

- Relative importance of sources (LT, λ , t), ... → map plasma populations (Ionosphere above Saturn's rings ?)
- Map fields, currents & waves (→ scattering of particles)
- B anomaly ? (or reexamine dynamo theory)
- Topology / Origin of radiation belts : Cosmic Rays interaction with rings ?
- Relative role of convection/corotation (intermediate between Earth and Jupiter ?)
- Dynamics of (outer) Magnetosphere (plasma blobs, plumes from Titan, Solar wind fluctuations ?)
- Direction-Finding of auroral radio source, nature of source, of rotation modulation ? Variations of radio rotation period ?
- Role of Titan ? Radio emission modulation by Dione ?

→ Complete mapping/exploration of plasma environment with MAPS package : CDA, INMS, MAG RPWS, CAPS, UVIS, MIMI (1st ENA imager embarked), +RSS

| CASSINI'S MAPS INSTRUMENTS | |
|-----------------------------------|---|
| Instrument | Objective |
| Cassini Plasma Spectrometer | Measures composition, density, velocity and temperature of ions and electrons |
| Cosmic Dust Analyzer | Measures flux, velocity, charge, mass and composition of dust and ice particles from 10^{-16} – 10^{-6} grams |
| Dual Technique Magnetometer | Measures the direction and strength of the magnetic field |
| Ion and Neutral Mass Spectrometer | Measures neutral species and low-energy ions |
| Magnetospheric Imaging Instrument | Images Saturn's magnetosphere using energetic neutral atoms, and measures the composition, charge state and energy distribution of energetic ions and electrons |
| Radio and Plasma Wave Science | Measures wave emissions as well as electron density and temperature |
| Radio Science Instrument | Measures the density of Saturn's ionosphere |
| Ultraviolet Imaging Spectrograph | Measures ultraviolet emissions to determine sources of plasma in Saturn's magnetosphere |

REFERENCES

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- ESA site on Cassini <http://www.esa.int/SPECIALS/Cassini-Huygens/>
- Details on historical background of Saturn's exploration <http://www.solarviews.com/eng/saturnbg.htm>